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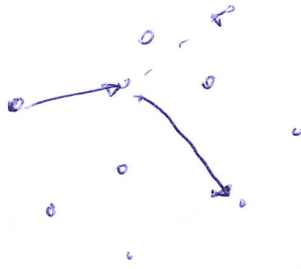
Introduction

P8_b

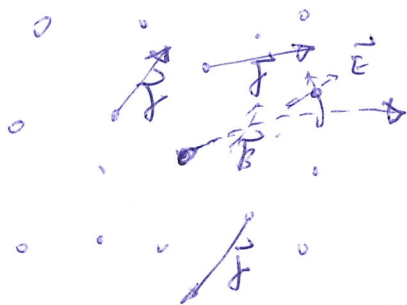
71.1 Definition of Plasma

- A plasma is a quasineutral gas of charged and neutral particles which exhibits collective behavior.

- * For a molecule, it moves undisturbed until it makes a "collision" with another.



- * For ~~charge~~ plasma, which has charged particles, fields effect the motion of other charged particles far away.



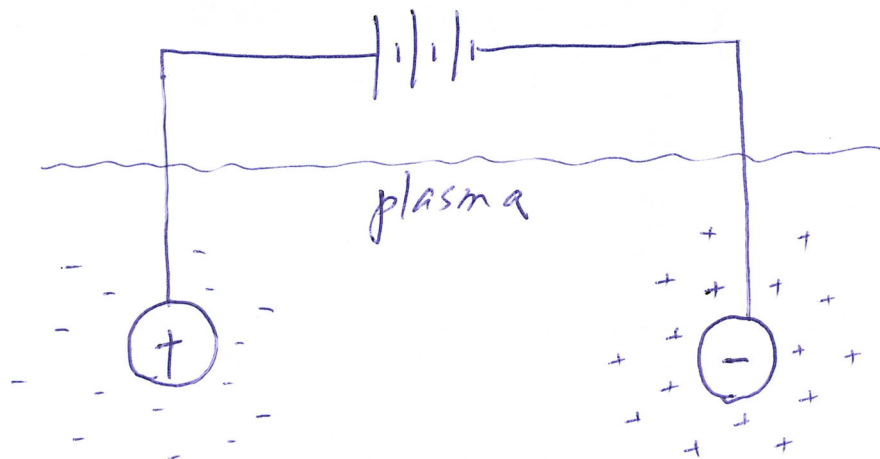
$$\vec{F} = q (\vec{E} + \vec{v} \times \vec{B})$$

- * "Collisionless" plasmas: the long-range electromagnetic forces are so much larger than the forces due to ordinary local collisions that the latter can be neglected altogether.

* "Collective behavior" means motion that depend not only on local conditions but on the state of the plasma in remote regions as well.

Q1.2. Debye Shielding

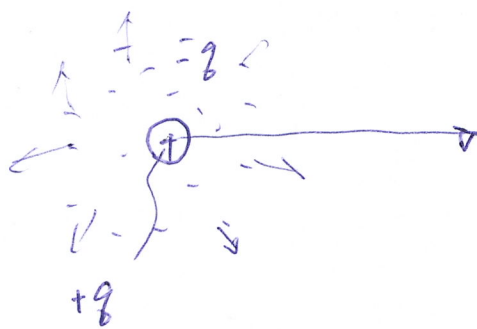
- A plasma is able to shield out electric potentials that are applied to it.



* If the plasma were "cold" ($T=0$) and there were no thermal motions, there would be just as many charges in the cloud as in the ball; the shielding would be perfect, and no electric field would be present in the body of the plasma outside of the clouds.

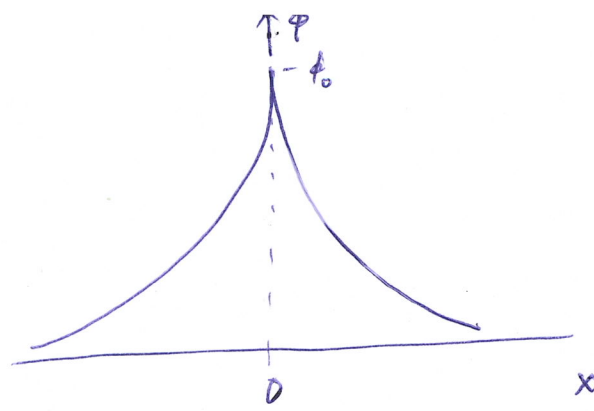
* If the temperature is "finite", those particles pro that are at the edge of the cloud, where the electric field is weak, have enough thermal energy to escape from the electrostatic potential well.

The "edge" of the cloud then occurs at the radius where the potential energy is approximately equal to the thermal energy kT of the particles, and the shielding is not complete. Potentials of the order of kT/e can ~~be~~ leak into the plasma and cause finite electric fields to exist here.



$$\left. \begin{aligned} V_+ &= \frac{kq}{r^2} \\ V_- &= \frac{k(L-q)}{r^2} \end{aligned} \right\} \Rightarrow V_+ + V_- = 0$$

$$V_- \neq \frac{k(L-q)}{r^2} \Rightarrow \underline{V_+ + V_- \neq 0}$$



Assuming $\frac{M}{m} \rightarrow \infty$ M : ion mass
 m : electron mass

\Rightarrow ions do not move but form a uniform background of positive charge.

Poisson's eq:

$$\nabla^2 \phi = -\frac{\rho}{\epsilon_0} \Rightarrow \epsilon_0 \frac{d^2 \phi}{dx^2} = -e(N_i - N_e) \text{ for } z=1$$

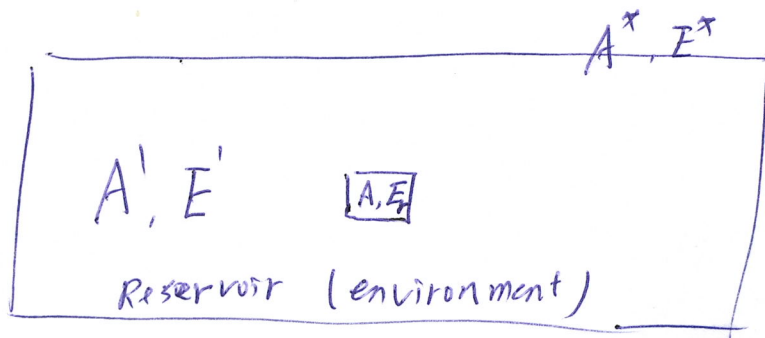
* Since ion doesn't move, $N_i = N_{\infty}$

* For electron, the electron distribution function:

$$f(u) = A \cdot \exp \left[-\frac{\frac{1}{2} m u^2 + e\phi}{k T_e} \right]$$

\rightarrow There are fewer particles at places where the potential energy is large, since not all particles have enough energy to get there.

* Boltzmann distribution (canonical distribution) p12



A^* - total system with energy E^*

A - small system with energy E_r at state r

A' - environment (reservoir) with energy E'

$$E^* = E_r + E', \quad E' \gg E_r$$

$\Omega(E) \rightarrow \#$ of states with energy E

\therefore the small system is at state r w/ energy E_r ,

\therefore ~~the small system is~~ $\Omega(E_r) = 1$

$$\Omega'(E') = \Omega'(E^* - E_r) = \# \text{ of states of the environment.}$$

\therefore the isolated system A^* is equally likely to be found in each one of its accessible states

$$\therefore P_r \propto \Omega'(E^* - E_r)$$

Note that $E^* \gg E_r$

$$\ln P_r \propto \ln [\Omega'(E^* - E_r)]$$

$$\approx \ln [\Omega(E^*)] - \frac{\partial \ln \Omega'}{\partial E'} E_r \equiv \ln [\Omega(E^*)] - \beta E_r$$

$$\Rightarrow P_r \propto \Omega'(E^* - E_r) \approx \Omega'(E^*) e^{-\beta E_r} \quad \beta \equiv \frac{1}{kT} \Rightarrow P_r = C e^{-\beta E_r}$$

To obtain $N_e(\phi)$

PL3

$$N_e(\phi) = \int_{-\infty}^{\infty} A e^{-\frac{(\frac{1}{2}mu^2 + q\phi)}{kT}} du$$

$q = -e$

$$= A e^{\frac{e\phi}{kT}} \int_{-\infty}^{\infty} e^{-\frac{1}{2}mu^2} du$$

$\stackrel{\text{I}}{=}$

$$I^2 = I \cdot I = \int_{-\infty}^{\infty} e^{-\frac{1}{2}mu^2} du \int_{-\infty}^{\infty} e^{-\frac{1}{2}mv^2} dv$$

$$= \int_{-\infty}^{\infty} \int_{-\infty}^{\infty} e^{-\frac{1}{2}m(u^2+v^2)} du dv$$

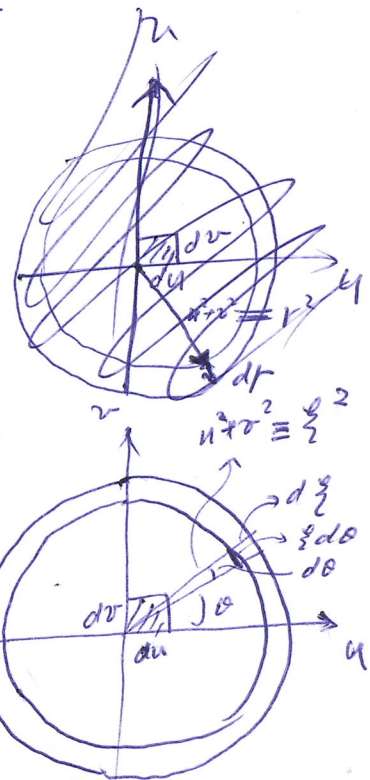
$$= \int_0^{\infty} \int_0^{2\pi} r e^{-\frac{1}{2}m r^2} dr d\theta$$

$$= \int_0^{2\pi} d\theta \int_0^{\infty} r e^{-\frac{1}{2}m r^2} dr$$

$$= 2\pi \int_0^{\infty} \frac{-dx}{m} e^x = -\frac{2\pi}{m} e^x \Big|_0^{\infty} = -\frac{2\pi}{m} e \Big|_0^{\infty}$$

$$= \frac{2\pi}{m} e^{-\frac{1}{2}m r^2} \Big|_0^{\infty} = \frac{2\pi}{m} \Rightarrow I = \sqrt{\frac{2\pi}{m}}$$

$$\Rightarrow N_e(\phi) = A \cdot e^{\frac{e\phi}{kT}} \cdot \sqrt{\frac{2\pi}{m}}$$



For $\phi \rightarrow 0$, i.e., $x \rightarrow \infty$, $n_e \rightarrow n_{\infty}$

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$$\Rightarrow n_e(0) = A \cdot \sqrt{\frac{2g}{m}} = n_{\infty}$$

$$\Rightarrow \underline{n_e = n_{\infty} e^{\frac{e\phi}{kT_e}}}$$

$$\begin{aligned} \Rightarrow \epsilon_0 \frac{d^2\phi}{dx^2} &= -e(n_i - n_e) \\ &= -e(n_{\infty} - n_{\infty} e^{\frac{e\phi}{kT_e}}) \\ &= e n_{\infty} \left[e^{\frac{e\phi}{kT_e}} - 1 \right] \end{aligned}$$

For $\frac{e\phi}{kT_e} \ll 1$, i.e. far away from the charge.

$$\epsilon_0 \frac{d^2\phi}{dx^2} \approx e n_{\infty} \left[1 + \left(\frac{e\phi}{kT_e}\right) + \frac{1}{2} \left(\frac{e\phi}{kT_e}\right)^2 + \dots - 1 \right]$$

$$\approx \frac{n_{\infty} e^2}{kT_e} \phi \quad \Rightarrow \quad \left(\frac{\epsilon_0 kT_e}{n e^2} \right) \frac{d^2\phi}{dx^2} = \phi$$

where $n \equiv n_{\infty}$

$$\Rightarrow \underline{\lambda_D \equiv \sqrt{\frac{\epsilon_0 kT_e}{n e^2}}} \quad \Rightarrow \quad \lambda_D^2 \frac{d^2\phi}{dx^2} = \phi$$

$$\Rightarrow \underline{\phi = \phi_0 \exp\left(-\frac{|x|}{\lambda_D}\right)}$$

λ_D is called the "Debye length" which is a measure of the shielding distance or thickness of the sheath.

* Useful forms:

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$$\lambda_D = 69 \left(\frac{T}{n} \right)^{1/2} \text{ (m)}, \quad T \text{ in keV}$$

$$= 7430 \left(\frac{kT}{n} \right)^{1/2} \text{ (m)}, \quad kT \text{ in "eV"}$$

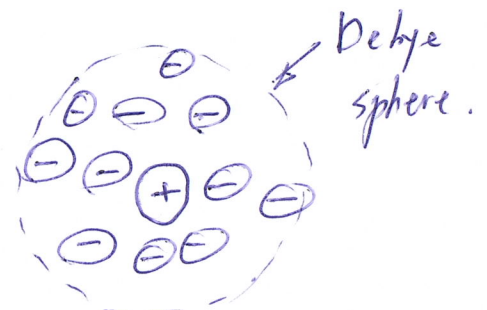
* "Quasineutrality": the dimensions of a system L are much ~~than~~ larger than λ_D , local charges are always shielded out in a distance short compared w/ L . \Rightarrow free of large electric potentials or fields.

$\Rightarrow n_i \simeq n_e \simeq n \rightarrow$ plasma density.

\rightarrow A criterion for an ionized gas to be a plasma is that it be dense enough that $\lambda_D \ll L$.

7.1.3 The plasma parameter.

* Debye shielding is valid only if there are enough ~~charged~~ particles in the charge cloud.



$$N_D = n \cdot \frac{4}{3} \pi \lambda_D^3$$

$$= \frac{4\pi}{3} n \cdot \left(\frac{\epsilon_0 k T_e}{n e^2} \right)^{3/2}$$

$$= 1.38 \times 10^6 \frac{T_e^{3/2}}{n^{1/2}}$$

T_e in Kelvin.
 n in $\frac{1}{m^3}$

"Collective behavior" requires.

$$N_D \gg 1$$

Q 1.4 Criteria for Plasma.

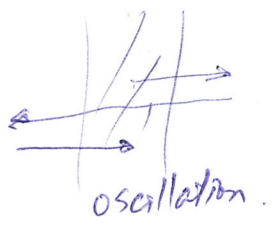
ω : frequency of typical plasma oscillations.

τ : mean time between collisions w/ ~~the~~ neutral atoms.

→ The weakly ionized gas in a jet exhaust, for example, ~~is~~ not ~~qualified~~ qualified as a plasma because the charged particles collide so frequently w/ neutral atoms that their motion is controlled by ordinary hydrodynamic forces rather than by electromagnetic forces.

$\omega \tau > 1$ is required to behave like a plasma

$\hookrightarrow 2\pi \frac{\tau}{T} > 1 \Rightarrow \tau > T \Rightarrow$ Not much collision w/ neutral gas within one oscillation.



Criteria for Plasma:

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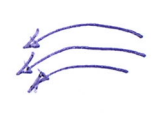
- 1. $\lambda_D \ll L$.
- 2. $N_D \gg 1$.
- 3. $\omega \tau > 1$.

Ch 2. Single-Particle Motions P18

* The first step is to understand how single particles behave in electric and magnetic fields.

* In this chapter, \vec{E} & \vec{B} are assumed to be prescribed and not affected by the charged particles

- * Uniform \vec{E} & \vec{B}
- $\vec{E} = 0, \vec{B} = \text{const.}$ - gyromotion
 - $\vec{E} = \text{const}, \vec{B} = \text{const.}$ - $\vec{E} \times \vec{B}$ drift
 - \vec{F} - gravitational field.

- * Non uniform \vec{B} ($\vec{E} = 0$)
- $\nabla B \perp \vec{B}$ - Grad-B drift.
 - Curved B - curvature drift 

- * Non uniform \vec{E} ($\vec{B} = \text{const}$)
(in space)

- * Time-varying \vec{E} (\vec{E}, \vec{B} uniform in space)
 $\vec{B}(t) = \text{const}(x)$

- * Time-varying \vec{B}
- Adiabatic invariants $\left\{ \begin{array}{l} \mu = \frac{mv_{\perp}^2}{2B} \\ J = \int_a^b v_{\parallel} ds \\ \mathcal{E} = \int \vec{B} \cdot d\vec{a} \end{array} \right.$

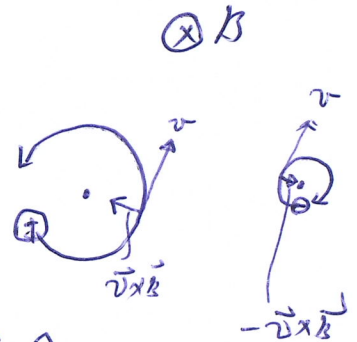
Q2. Uniform \vec{E} and \vec{B} .

Q2.1.1 $\vec{E} = 0$

- cyclotron gyration.

$$m \frac{d\vec{v}}{dt} = q(\vec{v} \times \vec{B})$$

Let $\vec{B} = B \hat{z}$



$$\begin{cases} m \dot{v}_x = qB v_y \\ m \dot{v}_y = -qB v_x \\ m \dot{v}_z = 0 \end{cases}$$

~~$v_z = 0$ $v_z(t=0) = 0$~~

$$m \ddot{v}_x = qB \dot{v}_y = -\frac{q^2 B^2}{m} v_x \Rightarrow \ddot{v}_x = -\left(\frac{qB}{m}\right)^2 v_x$$

$$m \ddot{v}_y = -qB \dot{v}_x = -\frac{q^2 B^2}{m} v_y \Rightarrow \ddot{v}_y = -\left(\frac{qB}{m}\right)^2 v_y$$

$$\omega_c = \frac{|q|B}{m} \quad \text{- cyclotron freq.}$$

$$\Rightarrow v_{x,y} = v_{\perp} \exp(\pm i\omega_c t + i\int dx/y) \quad \leftarrow \text{only the real part}$$

Choose the phase δ so that

$$\Rightarrow \begin{cases} v_x = v_{\perp} \exp(i\omega_c t) = \dot{x} \\ v_y = \frac{m}{qB} \dot{v}_x = \pm \frac{1}{\omega_c} \dot{v}_x = \pm i v_{\perp} e^{i\omega_c t} = \dot{y} \end{cases}$$

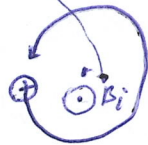
$$\Rightarrow \begin{cases} x = x_0 - r_L \frac{v_{\perp}}{\omega_c} e^{i\omega_c t} \\ y = y_0 \pm \frac{v_{\perp}}{\omega_c} e^{i\omega_c t} \end{cases} \Rightarrow r_L = \frac{v_{\perp}}{\omega_c} = \frac{m v_{\perp}}{|q|B}$$

$$\Rightarrow \text{real part.} \begin{cases} x = x_0 + r_L \sin(\omega_c t) \\ y = y_0 \pm r_L \cos(\omega_c t) \end{cases}$$

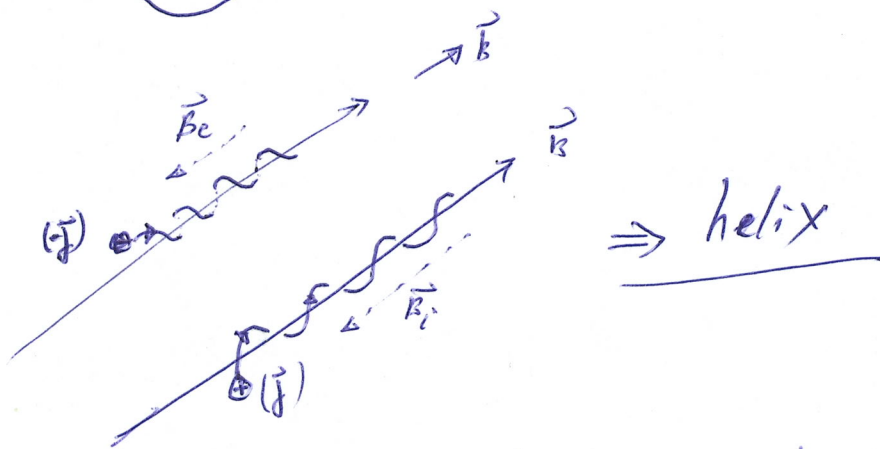
$$m \dot{v}_z = 0 \Rightarrow v_z = \text{const}, \quad z = v_z t. \quad \text{pro a}$$

$$\rightarrow \begin{cases} x = x_0 + r_L \sin(\omega_c t) \\ y = y_0 \pm r_L \cos(\omega_c t) \\ z = v_z t \end{cases} \quad \begin{cases} \omega_c = \frac{|q|B}{m} \\ r_L = \frac{v_\perp}{\omega_c} = \frac{m v_\perp}{|q|B} \end{cases}$$

generated B_i $\otimes B$



$v_z = 0$
 $\odot B_e$, generated B_e .

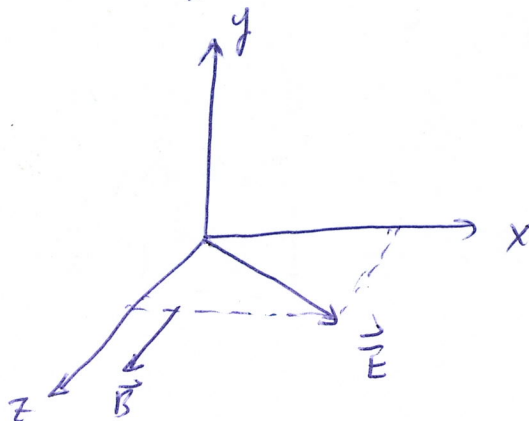


* The direction of the gyration is always such that the magnetic field generated by the charged particle is opposite to the externally imposed field.

q 2-1.2

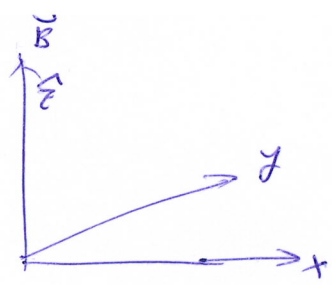
Finite \vec{E}

$$\vec{E} = \vec{E}_x + \vec{E}_z, \quad \vec{B} = \vec{B}_z$$



To derive the gyromotion again.

Let $\vec{B} = B \hat{z}$



$m \frac{d\vec{v}}{dt} = q(\vec{v} \times \vec{B})$ q can be either positive or negative.

$$\begin{cases} m \dot{v}_x = q B v_y & \Rightarrow \dot{v}_x = \frac{qB}{m} v_y \\ m \dot{v}_y = -q B v_x & \Rightarrow \dot{v}_y = -\frac{qB}{m} v_x \\ m \dot{v}_z = 0 & \Rightarrow v_z = 0, \quad z = v_z t + z_0 \end{cases}$$

$\ddot{v}_x = \frac{qB}{m} \dot{v}_y = \frac{qB}{m} (-\frac{qB}{m} v_x) = -(\frac{qB}{m})^2 v_x$, $\tilde{\omega} = \frac{qB}{m}$
 $\ddot{v}_y = -\frac{qB}{m} \dot{v}_x = -\frac{qB}{m} (\frac{qB}{m} v_y) = -(\frac{qB}{m})^2 v_y$
 $\tilde{\omega}$ can be positive or negative.

$\Rightarrow \begin{cases} \ddot{v}_x = -\tilde{\omega}^2 v_x \\ \ddot{v}_y = -\tilde{\omega}^2 v_y \end{cases} \Rightarrow \vec{v}_x = v_{\perp} e^{\pm i \tilde{\omega} t}$, $v_x = \text{Re} \{ \tilde{v}_x \}$, $v_y = \text{Re} \{ \tilde{v}_y \}$
 $\tilde{v}_y = \frac{m}{qB} \dot{\tilde{v}}_x = \frac{1}{\tilde{\omega}} v_{\perp} (\pm i \tilde{\omega}) e^{\pm i \tilde{\omega} t} = \pm i v_{\perp} e^{\pm i \tilde{\omega} t}$

$\Rightarrow v_x = \text{Re} \{ \tilde{v}_x \} = \text{Re} \{ v_{\perp} e^{\pm i \tilde{\omega} t} \} = \text{Re} \{ v_{\perp} (\cos \tilde{\omega} t \pm i \sin \tilde{\omega} t) \}$
 $= v_{\perp} \cos(\tilde{\omega} t)$

$v_y = \text{Re} \{ \tilde{v}_y \} = \text{Re} \{ \pm i v_{\perp} e^{\pm i \tilde{\omega} t} \} = \text{Re} \{ v_{\perp} (\pm i) (\cos \tilde{\omega} t \pm i \sin \tilde{\omega} t) \}$
 $= \text{Re} \{ v_{\perp} (-\sin \tilde{\omega} t \pm i \cos \tilde{\omega} t) \}$
 $= -v_{\perp} \sin(\tilde{\omega} t)$

~~for ion~~ define $\omega = \frac{|q|B}{m}$ for ion, $\omega = \tilde{\omega}$
 \uparrow always positive. for electron, $\omega = \frac{-|q|B}{m} = -\tilde{\omega}$

\Rightarrow For ion: $\begin{cases} v_x = v_{\perp} \cos \omega t \\ v_y = -v_{\perp} \sin \omega t \end{cases} \Rightarrow \begin{cases} v_x = v_{\perp} \cos(\omega t) \\ v_y = -v_{\perp} \sin(\omega t) \end{cases}$

For electron: $\begin{cases} v_x = v_{\perp} \cos(-\omega t) = v_{\perp} \cos \omega t \\ v_y = -v_{\perp} \sin(-\omega t) = v_{\perp} \sin \omega t \end{cases}$

$$x = \int v_x dt = \int v_{\perp} \cos \omega t dt = \frac{v_{\perp}}{\omega} \sin \omega t + x_0$$

P. 20 C

$$\equiv r_{\perp} \sin \omega t + x_0, \quad v_{\perp} \equiv \frac{v_x}{\omega}$$

$$y = \int v_y dt = \int \mp v_{\perp} \sin \omega t dt = \pm \frac{v_{\perp}}{\omega} \cos \omega t + y_0$$

$$\equiv \pm r_{\perp} \cos \omega t + y_0$$

$$\Rightarrow \begin{cases} x = r_{\perp} \sin(\omega t) + x_0 \\ y = \pm r_{\perp} \cos(\omega t) + y_0 \end{cases} \quad \begin{cases} v_x = v_{\perp} \cos(\omega t) \\ v_y = \mp v_{\perp} \sin(\omega t) \end{cases}$$

the upper sign is for ion.

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Let $x_0 = y_0 = 0$

$$\begin{cases} x = r_{\perp} \sin(\omega t) \\ y = \pm r_{\perp} \cos(\omega t) \end{cases}$$

for $t > 0$,

$$\begin{cases} x = 0 \\ y = \pm r_{\perp} \end{cases}$$

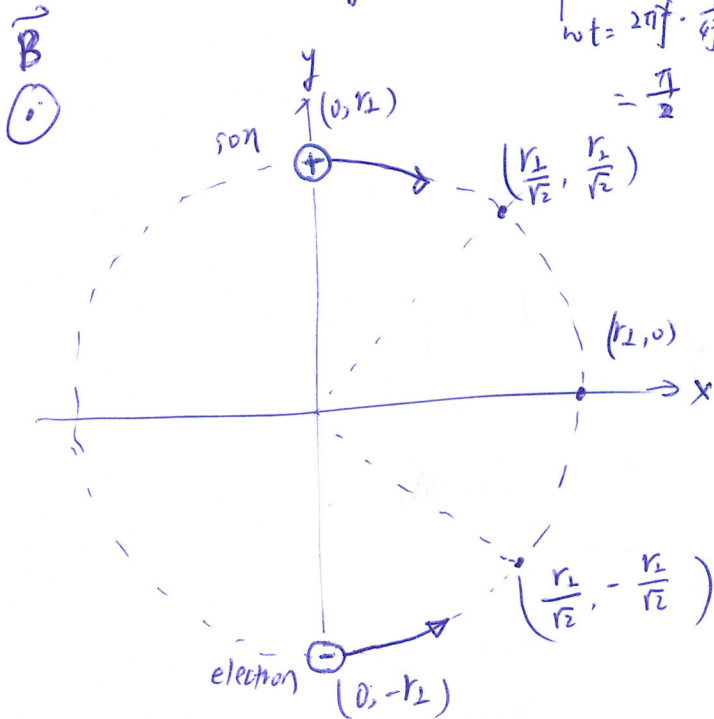
$$\begin{cases} t = \frac{T}{4} \\ = \frac{1}{4f} \\ \omega t = 2\pi f \cdot \frac{1}{4f} \\ = \frac{\pi}{2} \end{cases}$$

$$\begin{cases} x = r_{\perp} \\ y = 0 \end{cases}$$

$$\begin{cases} t = \frac{T}{8} \\ \omega t = \frac{\pi}{4} \end{cases}$$

$$x = \frac{r_{\perp}}{\sqrt{2}}$$

$$y = \pm \frac{r_{\perp}}{\sqrt{2}}$$



$$m \frac{d\vec{v}}{dt} = q (\vec{E} + \vec{v} \times \vec{B})$$

$$\vec{v} \times \vec{B} = \begin{vmatrix} \hat{x} & \hat{y} & \hat{z} \\ v_x & v_y & v_z \\ 0 & 0 & B \end{vmatrix} = \hat{x} (B \cdot v_y) + \hat{y} (-B \cdot v_x)$$

$$m \frac{dv_x}{dt} = q (E_x + B v_y)$$

$$m \frac{dv_y}{dt} = q (-B v_x)$$

$$m \frac{dv_z}{dt} = q E_z \Rightarrow v_z = \frac{q E_z}{m} t + v_{z0}$$

$$\frac{dv_x}{dt} = \frac{q}{m} E_x + \frac{qB}{m} v_y = \frac{q}{m} E_x \pm \omega_c v_y$$

$$\frac{dv_y}{dt} = -\frac{qB}{m} v_x = \mp \omega_c v_x$$

$$\omega_c = \frac{|q|B}{m}$$

$$\ddot{v}_x = \pm \omega_c \dot{v}_y = -\omega_c^2 v_x$$

$$-\ddot{v}_y = \mp \omega_c \dot{v}_x = \mp \omega_c \left(\frac{qB}{mB} E_x \pm \omega_c v_y \right) = -\omega_c^2 \left(v_y + \frac{E_x}{B} \right)$$

$$\text{let } v_y' = v_y + \frac{E_x}{B} \Rightarrow \dot{v}_y' = \dot{v}_y \text{ ; } \ddot{v}_y' = \ddot{v}_y$$

$$\Rightarrow \begin{cases} \ddot{v}_x = -\omega_c^2 v_x \\ \ddot{v}_y' = -\omega_c^2 v_y' \end{cases}$$

$$\Rightarrow v_x = v_{\perp} e^{i\omega_c t}$$

$$\Rightarrow v_y' = \pm i v_{\perp} e^{i\omega_c t}$$

$$\Rightarrow v_y = \pm i v_{\perp} e^{i\omega_c t} - \frac{E_x}{B}$$

a drift v_{ge} of the guiding center

$$\begin{cases} V_x = V_{\perp} e^{i\omega t} \\ V_y = \pm i V_{\perp} e^{i\omega t} - \frac{E_x}{B} \end{cases}$$

\vec{V}_{gy} - gyromotion \vec{V}_{gc} - guiding center drift of

Alternative way to derive:

$$\vec{v} \equiv \vec{v}_{gc} + \vec{v}_{gy}$$

$$m \frac{d\vec{v}}{dt} = m \frac{d(\vec{v}_{gc} + \vec{v}_{gy})}{dt} = m \frac{d\vec{v}_{gy}}{dt} = q [\vec{E} + (\vec{v}_{gc} + \vec{v}_{gy}) \times \vec{B}]$$

Take the time average of one cycle. $\langle m \frac{d\vec{v}_{gy}}{dt} \rangle = 0$
 $\langle \vec{v}_{gy} \rangle = 0$

$$\Rightarrow \vec{E} + \vec{v}_{gc} \times \vec{B} = 0 \quad \times \vec{B}$$

$$\vec{E} \times \vec{B} + (\vec{v}_{gc} \times \vec{B}) \times \vec{B} = 0$$

$$\begin{aligned} \Rightarrow \vec{E} \times \vec{B} &= \vec{B} \times (\vec{v}_{gc} \times \vec{B}) = (\vec{B} \cdot \vec{B}) \vec{v}_{gc} - (\vec{B} \cdot \vec{v}_{gc}) \vec{B} \\ &= v_{gc} B^2 - \vec{B} (\vec{v}_{gc} \cdot \vec{B}) \end{aligned}$$

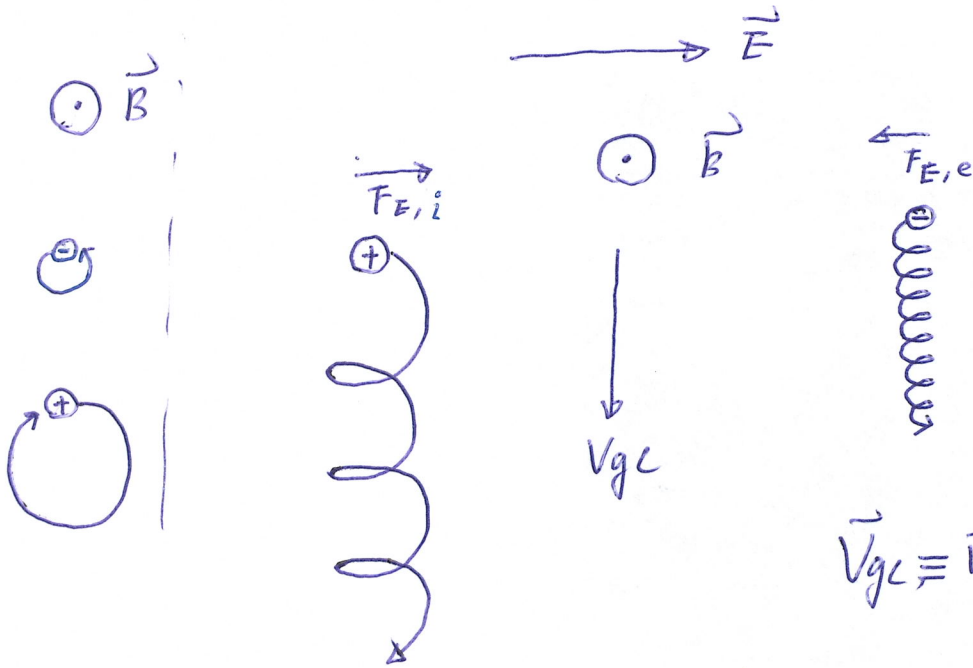
Note that $\vec{B} = B \hat{z}$, i.e. $\vec{E} \times \vec{B} \perp \hat{z}$

$$\text{let } \vec{v}_{gc} = v_{gc,z} \hat{z} + \vec{v}_{gc,\perp}$$

$$\Rightarrow \underbrace{\vec{E} \times \vec{B}}_{\perp \hat{z}} = \cancel{v_{gc,z} B^2 \hat{z}} + \vec{v}_{gc,\perp} B^2 - \cancel{B^2 v_{gc,z} \hat{z}}$$

$$\Rightarrow \vec{v}_{gc,\perp} = \frac{\vec{E} \times \vec{B}}{B^2} \equiv \vec{v}_E ; v_E = \frac{E(V/m)}{B(T)} \frac{m}{s}$$

V_E is independent of q, m, v_{\perp} !!



$\vec{v}_{gc} \equiv \vec{v}_E$ is called the $\vec{E} \times \vec{B}$ drift

$$\omega_c = \frac{|q|B}{m}$$

$$r_L = \frac{v_{\perp}}{\omega_c} = \frac{mv_{\perp}}{|q|B}$$

$\left. \begin{matrix} \omega_i < \omega_e \\ r_i > r_e \end{matrix} \right\} \rightarrow \begin{matrix} \wedge \\ \text{Two effects} \end{matrix}$ canceled out

§ 2.1.3 Gravitational Field

$$\vec{F}_{em} = q(\vec{E} + \vec{v} \times \vec{B}) = q\vec{E} + q\vec{v} \times \vec{B} = \vec{F}_E + \vec{F}_m$$

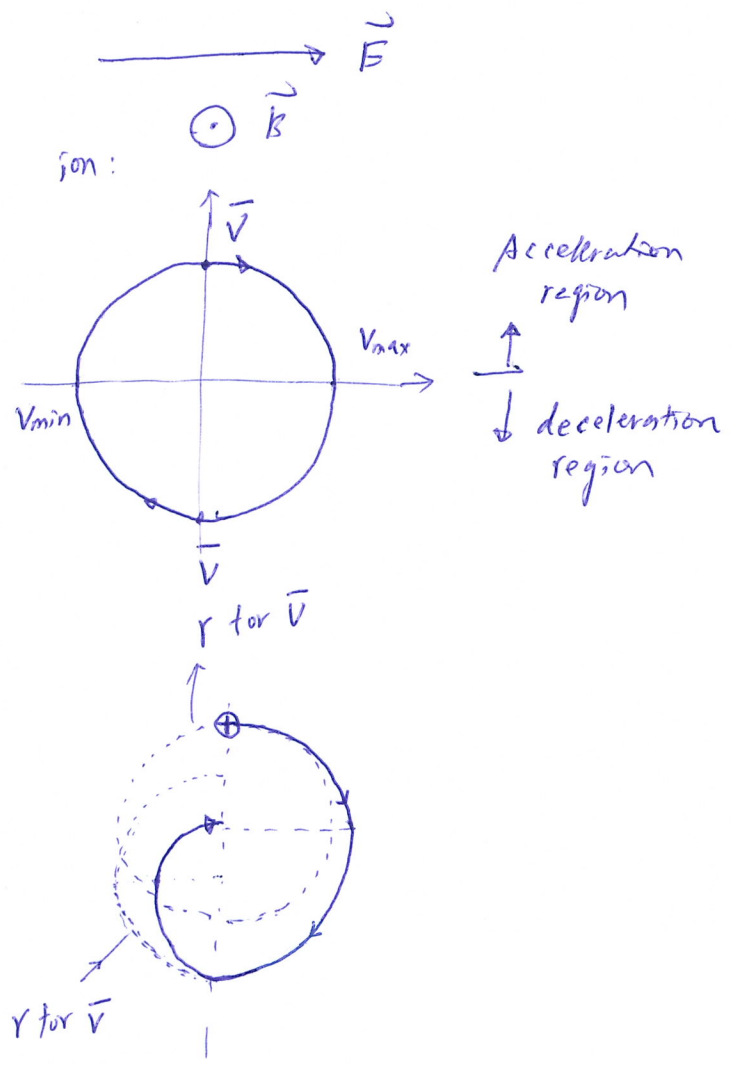
↑
electrical force

\vec{F}_E can be any kind of force

i.e. $m \frac{d\vec{v}}{dt} = \vec{F} + q\vec{v} \times \vec{B} \iff m \frac{d\vec{v}}{dt} = q\vec{E} + q\vec{v} \times \vec{B}$

$$\Rightarrow \vec{v}_f = \frac{(\vec{F}/q) \times \vec{B}}{B^2} = \frac{1}{q} \frac{\vec{F} \times \vec{B}}{B^2}$$

$$\vec{v}_E = \frac{\vec{E} \times \vec{B}}{B^2}$$

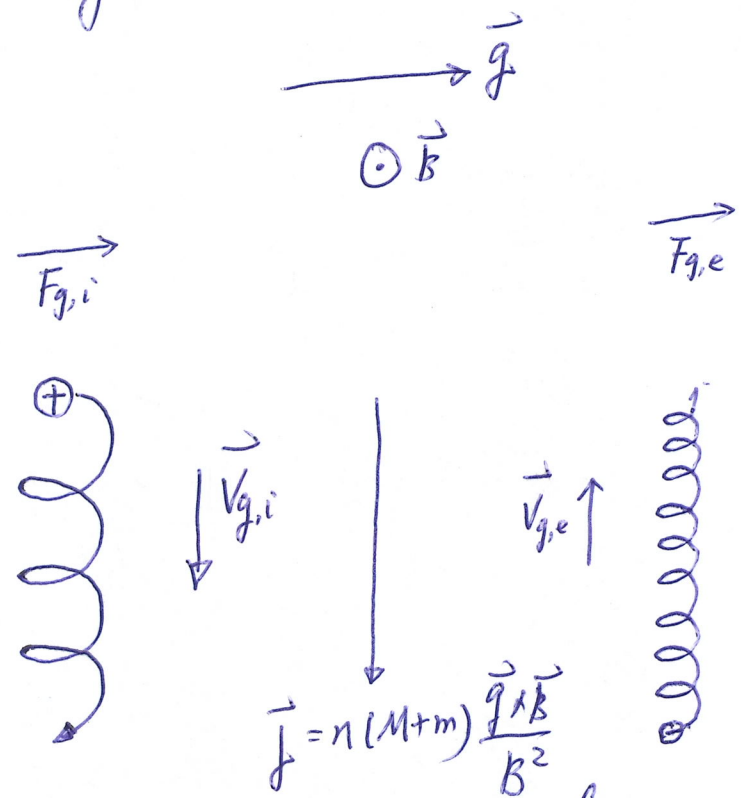


If \vec{F}_g is the force of gravity; i.e.,

$$\vec{F}_g = m \vec{g}$$

$$\Rightarrow \vec{V}_g = \frac{m}{q} \frac{\vec{g} \times \vec{B}}{B^2} \quad \text{gravitational drift}$$

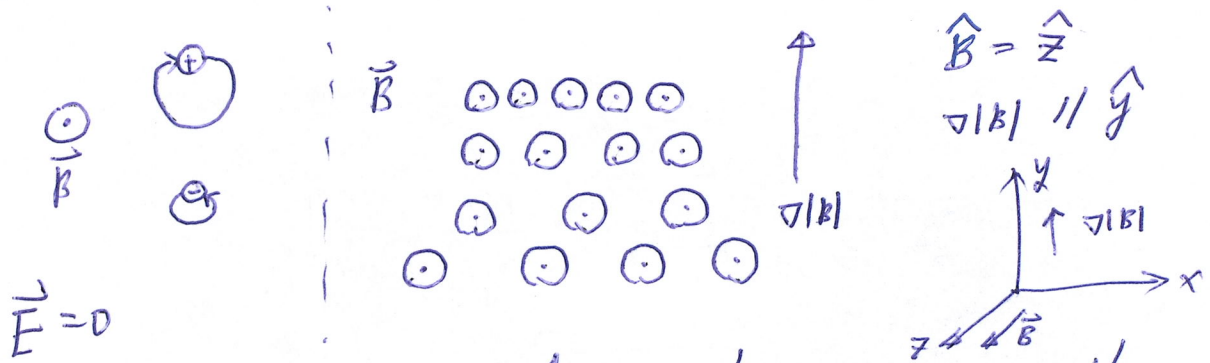
* drift \vec{V}_g changes sign with the particles charge.



* Under a gravitational force, ions and electrons drift in opposite directions, so there is a net current density in the plasma.

* \vec{V}_g is usually negligible. (homework?)

§ 2.2 Nonuniform \vec{B} ($\vec{E}=0$)



* When we introduce inhomogeneity, the problem becomes too complicated to solve exactly.

"Orbit Theory", which is customary to expand on the small ratio r_L/L , is used.

L is the scale length of the inhomogeneity.

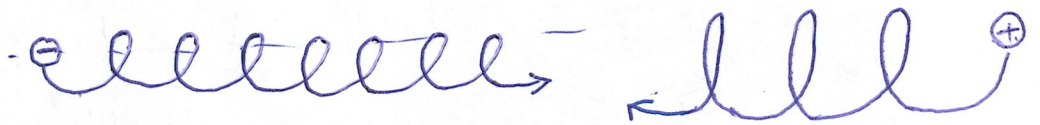
$$L = \left(\frac{1}{B} \nabla B \right)^{-1}$$

§ 2.2.1 $\nabla B \perp \vec{B}$: Grad-B drift

→ Lines of force are straight, but their density increases.

$$r_L = \frac{m v_{\perp}}{|B|} \propto \frac{1}{B}$$

\vec{B} large
 \odot
 \vec{B} small

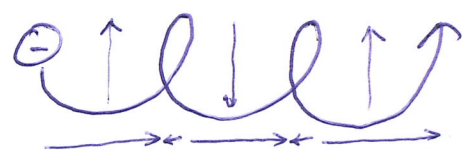


$$\Rightarrow v_{\text{drift}} \propto v_{\perp}, \frac{r_L}{L}$$

$$\vec{F} = q \vec{v} \times \vec{B}$$

average

* $\overline{F_x} = 0$ since the particle spends as much time moving up as down.



* To calculate $\overline{F_y}$, use the "undisturbed orbit" of the particle to find the average.

i.e. pure gyromotion.

$$\begin{cases} V_x = V_{\perp} e^{i\omega_c t} \\ V_y = \pm i V_{\perp} e^{i\omega_c t} \end{cases} \Rightarrow \begin{matrix} \text{real} \\ \text{part} \end{matrix} \quad \begin{cases} V_x = V_{\perp} \cos(\omega_c t) \\ V_y = \mp V_{\perp} \sin(\omega_c t) \end{cases}$$

$$\vec{F} = q \vec{v} \times \vec{B} = q \begin{vmatrix} \hat{x} & \hat{y} & \hat{z} \\ V_x & V_y & 0 \\ 0 & 0 & B \end{vmatrix} = \hat{x} (q V_y B) + \hat{y} (q V_x B)$$

Consider $\overline{F_y}$ only.

$\because \nabla B \parallel \hat{y}$

$$F_y = -q V_x B = -q V_{\perp} \cos(\omega_c t) \cdot B(y)$$

$$\vec{B} = \vec{B}_0 + (\vec{r} \cdot \nabla) \vec{B} + \dots$$

$$B_z(y) = B_0 + y \cdot \frac{\partial B}{\partial y} + \dots \quad \begin{aligned} y &= y_0 \pm r_{\perp} \cos(\omega_c t) \\ &\equiv \pm r_{\perp} \cos(\omega_c t) \end{aligned}$$

$$\Rightarrow F_y = -q V_{\perp} \cos(\omega_c t) \left[B_0 \pm \underbrace{r_{\perp} \cos(\omega_c t)}_y \frac{\partial B}{\partial y} \right]$$

Note that $\frac{r_L}{L} \ll 1$, L is the scale length ^(p2)

i.e. $L \equiv \left(\frac{1}{B} \frac{\partial B}{\partial y} \right)^{-1} \gg r_L$

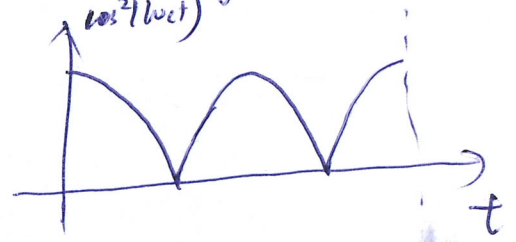
$$B(y) = B_0 + y \cdot \frac{\partial B}{\partial y} + \dots$$

$$= B_0 + y \cdot \frac{B_0}{L} + \dots$$

$$= B_0 \left[1 + \frac{y}{L} + \dots \right]$$

$y \sim r_L$, $\therefore \frac{y}{L} \sim \frac{r_L}{L} \ll 1 \rightarrow$ keep the 1st term

$$F_y = -g v_{\perp} B_0 \cos(\omega ct) + g v_{\perp} r_L \frac{\partial B}{\partial y} \cos^2(\omega ct)$$

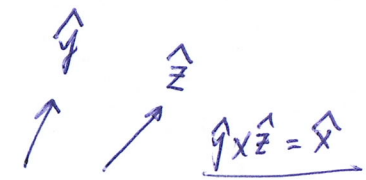


$$\overline{\cos(\omega ct)} = \frac{1}{T} \int_0^T \cos(\omega ct) dt = 0$$

$$\overline{\cos^2(\omega ct)} = \frac{1}{T} \int_0^T \cos^2(\omega ct) dt = \frac{1}{2}$$

$$\overline{F_y} = + \frac{1}{2} g v_{\perp} r_L \left(\frac{\partial B}{\partial y} \right)$$

"Gravitational Drift" $\vec{v}_{gc} = \frac{1}{g} \frac{\vec{F} \times \vec{B}}{B^2}$



$$\vec{v}_{gc} = \frac{1}{g} \frac{\overline{F_y} \cdot \vec{B}}{B^2} \hat{x} = \frac{1}{g} \frac{\overline{F_y}}{B} \hat{x} = + \frac{1}{2} \frac{v_{\perp} r_L}{B} \left(\frac{\partial B}{\partial y} \right) \hat{x}$$

$$\frac{\partial B}{\partial y} \rightarrow \nabla B$$

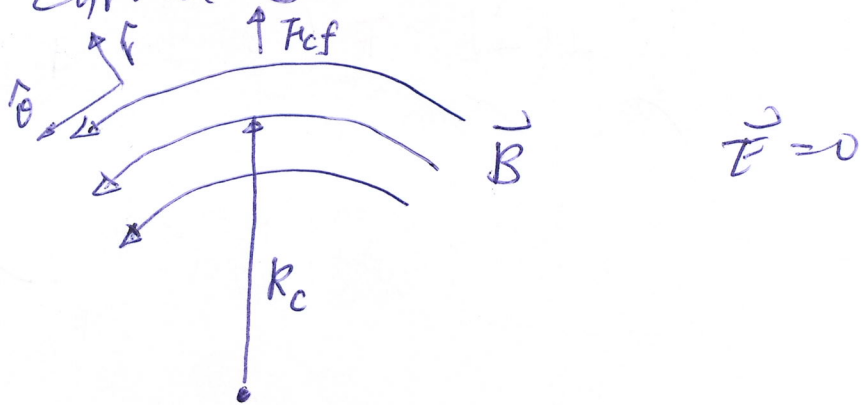
$$\hat{x} = \hat{y} \times \hat{z} \rightarrow \hat{B} \times \nabla B$$

unit vector.

$$\Rightarrow \vec{V}_{\nabla B} = \pm \frac{1}{2} v_{\perp} r_{\perp} \frac{\hat{B} \times \nabla B}{B^2} \quad \text{grad-B drift}$$

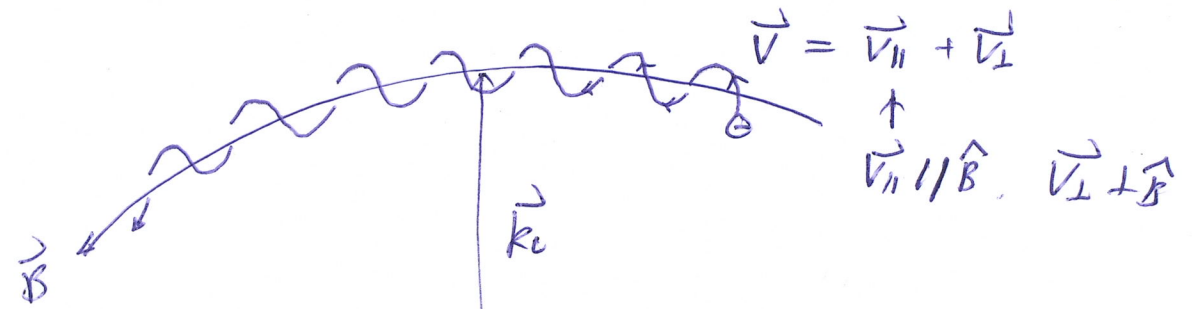
* It is in opposite directions for ions and electrons and causes a current transverse to \vec{B} .

§ 2.2.2 Curved B: Curvature Drift. ($\vec{E} = 0$)



Assume

- * Lines of force to be curved with a constant radius of curvature R_c , and $|B|$ to be constant.
- * The field does not obey Maxwell's eq. on a vacuum.
- * A guiding center drift arises from the "centrifugal force" felt by the particles as they move along the field lines in their thermal motion.



$$\vec{F}_{cf} = \frac{m v_{||}^2}{R_c} \quad \hat{r} = m v_{||}^2 \frac{\vec{R}_c}{R_c^2}$$

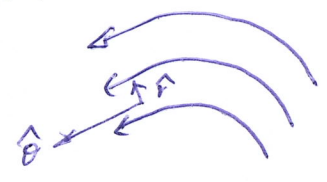
$$\vec{V}_{ge} = \frac{1}{B} \frac{\vec{F} \times \vec{B}}{B^2} \Rightarrow \vec{V}_R = \frac{1}{B} \frac{\vec{F}_{cf} \times \vec{B}}{B^2} = \frac{m v_{||}^2}{2 B^2} \frac{\vec{R}_c \times \vec{B}}{R_c^2}$$

Curvature drift. ^{*}

* Consider the decrease of $|B|$ w/ radius.

- Note that $\nabla \times \vec{B} = 0$ in vacuum.

- In cylindrical coordinates of



$\nabla \times \vec{B} \parallel \hat{z}$

$$\begin{aligned} \because \vec{B} = B \hat{\theta}, \quad \nabla B \parallel \hat{r} \\ \therefore \nabla \times \vec{B} = \frac{1}{r} \begin{vmatrix} \hat{r} & r\hat{\theta} & \hat{z} \\ \frac{\partial}{\partial r} & \frac{\partial}{\partial \theta} & \frac{\partial}{\partial z} \\ B_r & rB_{\theta} & B_z \end{vmatrix} = \frac{1}{r} \begin{vmatrix} r & r\hat{\theta} & \hat{z} \\ \frac{\partial}{\partial r} & 0 & 0 \\ 0 & rB_{\theta} & 0 \end{vmatrix} \\ = \hat{z} \frac{1}{r} \frac{\partial (r B_{\theta})}{\partial r} \end{aligned}$$

$\nabla \times \vec{B} = 0$

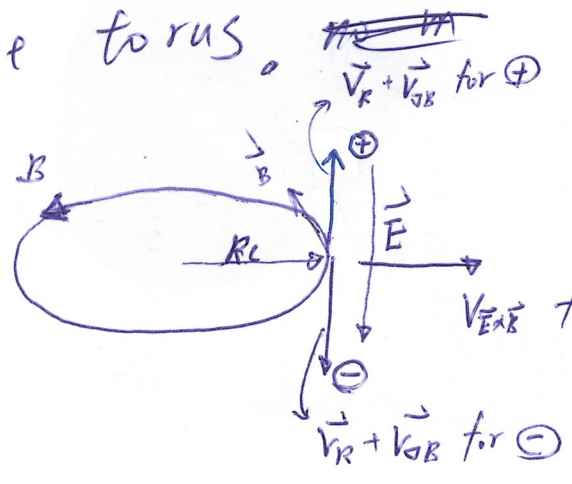
$$(\nabla \times \vec{B})_z = \frac{1}{r} \frac{\partial r B_{\theta}}{\partial r} = 0 \Rightarrow B_{\theta} \propto \frac{1}{r}$$

$$|B| \propto \frac{1}{R_c} \quad \frac{\nabla |B|}{|B|} = \frac{1}{R_c} \cdot \frac{\partial \frac{1}{r}}{\partial r} \Big|_{R_c} \hat{r} = -\frac{R_c}{R_c^2} \hat{r} = -\frac{\vec{R}_c}{R_c^2}$$

$$\begin{aligned}
 V_{\nabla B} &= \pm \frac{1}{2} v_{\perp} r_L \frac{\vec{B} \times \nabla B}{B^2} \frac{|B|}{|B|} - \frac{\vec{R}_c}{R_c^2} \\
 &= \pm \frac{1}{2} v_{\perp} r_L \frac{|B|}{B^2} \vec{B} \times \frac{\nabla B}{|B|} \\
 &= \mp \frac{1}{2} v_{\perp} r_L \frac{1}{B} \vec{B} \times \frac{\vec{R}_c}{R_c^2} \quad r_L = \frac{v_{\perp}}{\omega_c} \\
 &= \pm \frac{1}{2} \frac{v_{\perp}^2}{\omega_c} \frac{\vec{R}_c \times \vec{B}}{R_c^2 B} \quad \omega_c = \frac{|B|}{m} \\
 &= \pm \frac{1}{2} \frac{v_{\perp}^2}{|B| B / m} \frac{\vec{R}_c \times \vec{B}}{R_c^2 B} \\
 &= \pm \frac{1}{2} \frac{m}{B} v_{\perp}^2 \frac{\vec{R}_c \times \vec{B}}{R_c^2 B^2}
 \end{aligned}$$

$$\vec{V}_R + \vec{V}_{\nabla B} = \frac{m}{q} \frac{\vec{R}_c \times \vec{B}}{R_c^2 B^2} \left(v_{\parallel}^2 + \frac{1}{2} v_{\perp}^2 \right)$$

* If one bends a magnetic field into a torus for the purpose of containing a thermonuclear plasma, the particle will drift out of the torus.



* Curve + Grad B drifts cause charge separation in a ~~Tokamak~~ torus (Tokamak)

7 2.2.3. $\nabla \parallel \vec{B}$: Magnetic Mirrors

$\vec{B} = B(z) \hat{z}$

Axisymmetric:

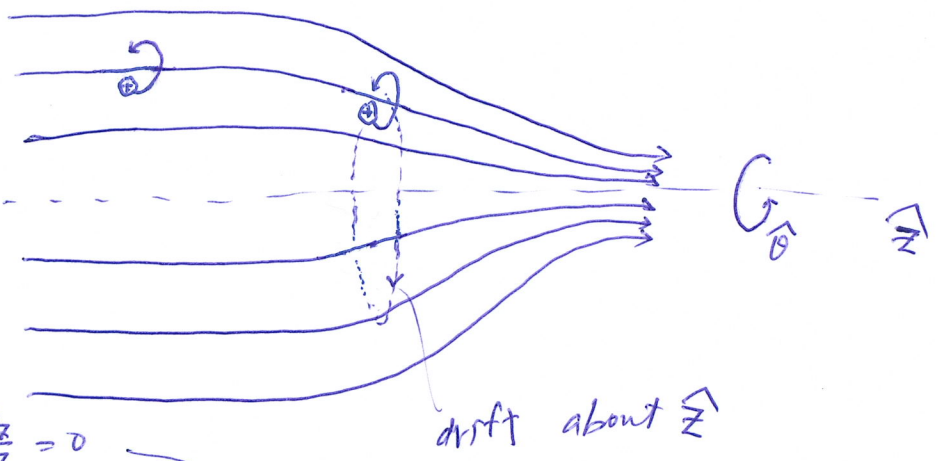
$B_\theta = 0, \frac{\partial}{\partial \theta} = 0$

$B_r \neq 0$

$\nabla \cdot \vec{B} = 0$

$\frac{1}{r} \frac{\partial}{\partial r}(r B_r) + \frac{1}{r} \frac{\partial B_\theta}{\partial \theta} + \frac{\partial B_z}{\partial z} = 0$

$\left(\Rightarrow \frac{1}{r} \frac{\partial}{\partial r}(r B_r) = -\frac{\partial B_z}{\partial z} \neq 0 \right)$



$\frac{1}{r} \frac{\partial}{\partial r}(r B_r) + \frac{\partial B_z}{\partial z} = 0$

If \textcircled{a} $r=0$, $\frac{\partial B_z}{\partial z}$ is given & does not vary much w/ r (weak funct. of r)

$\Rightarrow \frac{\partial}{\partial r}(r B_r) = -r \frac{\partial B_z}{\partial z}$

$\Rightarrow r B_r = - \int_0^r r \frac{\partial B_z}{\partial z} dr \approx - \left[\frac{\partial B_z}{\partial z} \right]_{r=0} \int_0^r r dr = -\frac{1}{2} r^2 \left[\frac{\partial B_z}{\partial z} \right]_{r=0}$

$\Rightarrow B_r = -\frac{1}{2} r \left[\frac{\partial B_z}{\partial z} \right]_{r=0}$

$\Rightarrow \vec{B}$ has component both in \hat{r} & \hat{z}
 $\nabla |B|$ is in \hat{z} .
 $\Rightarrow \nabla |B| \times \vec{B}$ is in θ direction
 Alternatively, $\nabla |B|$ has \hat{r} component,
 \vec{B} is in $r-z$ plane.
 $\nabla |B| \times \vec{B}$ is in θ direction

The variation of $|B|$ w/ r causes a grad-B drift of guiding centers about the axis of symmetry (\hat{z}).

* $\therefore \frac{\partial B}{\partial \theta} = 0$

\therefore NO radial grad-B drift.

Lorentz force: $F = q(\vec{v} \times \vec{B}) = q \begin{vmatrix} \hat{r} & \hat{\theta} & \hat{z} \\ v_r & v_\theta & v_z \\ B_r & B_\theta & B_z \end{vmatrix}$ p32

$$\begin{cases} F_r = q (v_\theta B_z - v_z B_\theta) \\ F_\theta = q (-v_r B_z + v_z B_r) \\ F_z = q (v_r B_\theta - v_\theta B_r) \end{cases}$$

① + ② : usual Larmor gyration.

③ : $(v_z B_r) \rightarrow \neq 0$ ④ axis, $\therefore r \neq 0$
 $\neq 0$: a azimuthal force causes a drift in the radial direction.

④ : $F_z = -q v_\theta B_r = \frac{1}{2} q v_\theta r \left[\frac{\partial B_z}{\partial z} \right]$ ← interesting.

* - Average over one gyration.

- Consider a particle whose guiding center lies on the axis

$\Rightarrow v_\theta$ is a const.

$v_\theta = \mp v_\perp$ depend on q .

$r = r_L$ ← Larmor radius.

$$\frac{1}{2} q v_\theta r \left[\frac{\partial B_z}{\partial z} \right]$$

$$r_L = \frac{v_\perp}{\omega_c} = \frac{m v_\perp}{|q| B}$$

$$\omega_c = \frac{|q| B}{m}$$

$$\begin{aligned} \bar{F}_z &= -\frac{1}{2} |q| v_\perp r_L \left[\frac{\partial B_z}{\partial z} \right] = -\frac{1}{2} |q| \frac{m v_\perp^2}{|q| B} \left[\frac{\partial B_z}{\partial z} \right] \\ &= -\frac{1}{2} \frac{m v_\perp^2}{B} \left[\frac{\partial B_z}{\partial z} \right] = -\mu \left[\frac{\partial B_z}{\partial z} \right] \quad \text{where} \end{aligned}$$

$$\mu \equiv \frac{1}{2} \frac{m v_\perp^2}{B} \quad \text{— magnetic moment.}$$

In general, $\hat{\uparrow}$ diamagnetic particle:
force on a

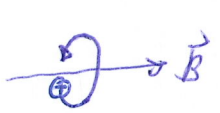
29
p33

$$\vec{F}_{||} = -\mu \frac{\partial B}{\partial s} = -\mu \nabla_{||} B$$

$d\vec{s}$: a line element along \vec{B}

* Note that the magnetic moment μ is the same as the magnetic moment of a current loop w/ area A & current I :

$$\mu = I \cdot A \quad \omega = 2\pi f = \frac{2\pi}{T}$$



$$I = \frac{q}{T} = q \cdot \frac{\omega_c}{2\pi}, \quad A = \pi r_L^2 = \pi \frac{v_{\perp}^2}{\omega_c^2}$$

$$\mu = q \cdot \frac{v_{\perp}^2}{2\pi} \cdot \pi \frac{v_{\perp}^2}{\omega_c^2} = \frac{q}{2} \frac{v_{\perp}^2}{\omega_c}$$

$\omega_c = \frac{|q|B}{m}$
 $f = \frac{qB}{m}$

$$= \frac{q}{2} \frac{m v_{\perp}^2}{qB} = \frac{1}{2} \frac{m v_{\perp}^2}{B}$$

$\rightarrow B \rightarrow$ stronger/weaker \rightarrow Larmor radius r_L changes

$\Rightarrow \mu$ remains invariant.

$$F_{||} = m \frac{dv_{||}}{dt} = -\mu \frac{\partial B}{\partial s} \quad \times v_{||}$$

$$m v_{||} \frac{dv_{||}}{dt} = \frac{d}{dt} \left(\frac{1}{2} m v_{||}^2 \right) = -\mu \frac{\partial B}{\partial s} \cdot v_{||} = -\mu \frac{\partial B}{\partial s} \frac{ds}{dt} = -\mu \frac{dB}{dt}$$

$$\frac{dB}{dt} = \frac{\partial B}{\partial t} + \frac{\partial B}{\partial s} \frac{ds}{dt} \quad \leftarrow \text{Variation of } B \text{ seen by the particle}$$

B doesn't change w/ time $\Rightarrow \frac{\partial B}{\partial t} = 0$

Note that the particle's energy is conserved \leftarrow why? (HW)

$$\frac{d}{dt} \left(\frac{1}{2} m v_{||}^2 + \frac{1}{2} m v_{\perp}^2 \right) = \frac{d}{dt} \left(\frac{1}{2} m v_{||}^2 + \mu B \right) = 0$$

$\therefore \mu = \frac{1}{2} \frac{m v_{\perp}^2}{B}$

$$\frac{d}{dt} \left(\frac{1}{2} m v_{\perp}^2 + \mu B \right) = 0 \Rightarrow \frac{d}{dt} \left(\frac{1}{2} m v_{\perp}^2 \right) + \frac{d}{dt} (\mu B) = 0$$

$$\Rightarrow -\mu \frac{dB}{dt} + \frac{d}{dt} (\mu B) = 0 \Rightarrow \mu \frac{d\mu}{dt} = 0$$

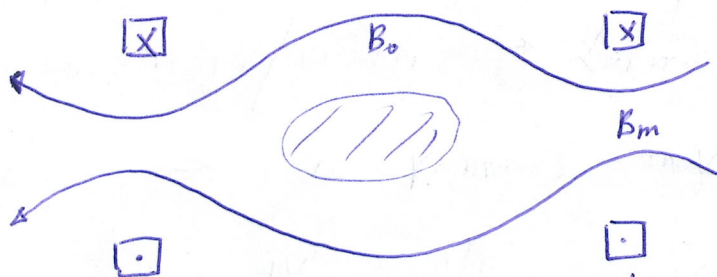
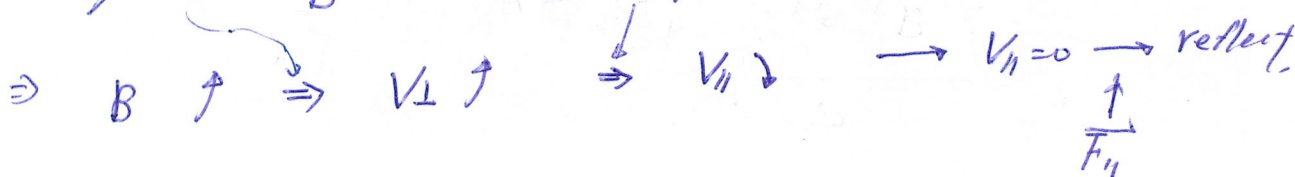
$$\mu \frac{dB}{dt} + B \frac{d\mu}{dt}$$

$$\Rightarrow \frac{d\mu}{dt} = 0$$

* The invariance of μ is the basis for one of the primary schemes for plasma confinement: the magnetic mirror

$$\mu = \frac{1}{2} \frac{m v_{\perp}^2}{B}$$

energy conservation.



throat of the mirror

* The trapping is not perfect:

- ① $v_{\perp} = 0$, no magnetic moment, \Rightarrow no force along \vec{B} .
- ② small $\frac{v_{\perp}}{v_{\parallel}}$ ③ mid plane ($B = B_0$) may escape if B_m is not enough.

For a particle:

① mid plane: $V_{\perp} = V_{\perp 0}$, $V_{\parallel} = V_{\parallel 0}$, $B = B_0$

② turning point: $V_{\perp} = V_{\perp}'$, $V_{\parallel} = V_{\parallel}'$, $B = B'$

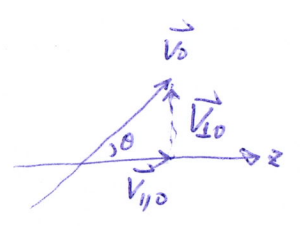
$$\mu = \frac{1}{2} \frac{m V_{\perp 0}^2}{B_0} = \frac{1}{2} \frac{m V_{\perp}'^2}{B'} \Rightarrow \frac{V_{\perp 0}^2}{B_0} = \frac{V_{\perp}'^2}{B'}$$

Energy conservation:

$$V_{\parallel}^2 + V_{\perp}'^2 \equiv V_{\perp}'^2 = V_{\parallel 0}^2 + V_{\perp 0}^2 \equiv V_0^2$$

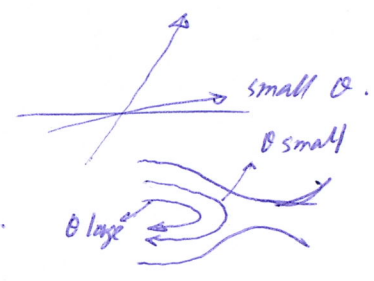
$V_{\parallel}^2 = 0$ @ turning point

$$\Rightarrow \frac{B_0}{B'} = \frac{V_{\perp 0}^2}{V_{\perp}'^2} = \frac{V_{\perp 0}^2}{V_0^2} = \sin^2 \theta$$



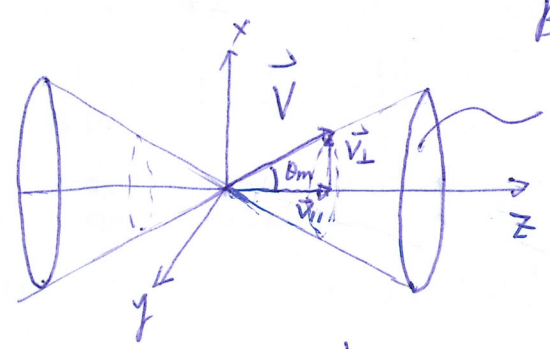
* θ is the pitch angle of the orbit in the weak-field region.

- For particle w/ smaller θ \rightarrow mirror in a region of higher B .



- If $B' > B_m \rightarrow$ the particle does not mirror. i.e., $V_{\parallel}' \neq 0$

The smallest θ : $\frac{B_0}{B_m} = \sin^2 \theta_m \equiv \frac{1}{R_m} \leftarrow$ mirror ratio



loss cone. \rightarrow particle in this region can not be confined.

Velocity space.

* loss cone is independent of β , m . p31
→ electrons & ions are equally confined.
w/o collisions

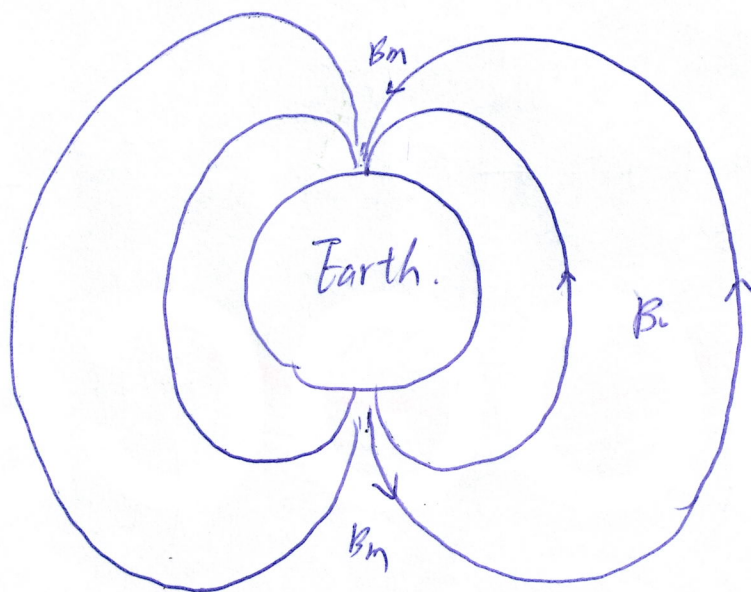
* Mirror confined plasma is NEVER isotropic.

* w/ collisions, particles are lost when they change their pitch angle in a collision and are scattered into the loss cone.

* Electrons have larger collision freq. than ions
⇒ electrons are lost more easily.

* 1st mirror motion was proposed by Enrico Fermi

* Another example: Van Allen belts.

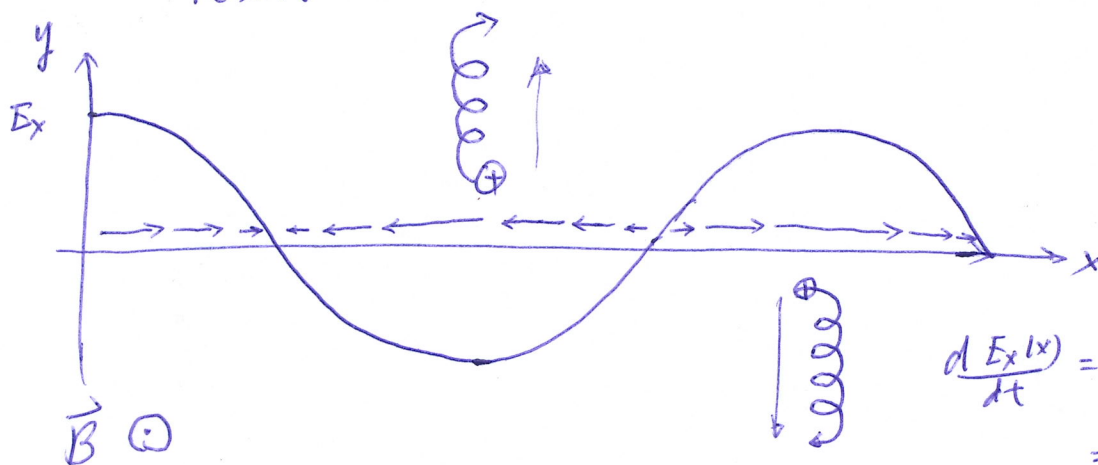


7. 2.3 Nonuniform E field.

$B = \text{uniform} = B \hat{z}$

$\vec{E} = E_0 \cos(kx) \hat{x} \quad \lambda = \frac{2\pi}{k}$

↳ result of a sinusoidal distribution of charges.



$$\frac{dE_x(x)}{dt} = \frac{\partial E_x(x)}{\partial t} + \frac{\partial E_x}{\partial x} \frac{dx}{dt}$$

$$= \frac{\partial E_x}{\partial x} \cdot \frac{dx}{dt} \neq 0$$

$$m \frac{d\vec{v}}{dt} = q [\vec{E}(x) + \vec{v} \times \vec{B}]$$

$$\begin{cases} \dot{v}_x = \frac{\partial B}{\partial m} v_y + \frac{q}{m} E_x(x) & \Rightarrow \ddot{v}_x = \frac{\partial B}{\partial m} \dot{v}_y + \frac{q}{m} \dot{E}_x(x) \\ \dot{v}_y = -\frac{\partial B}{\partial m} v_x & \Rightarrow \ddot{v}_y = -\frac{\partial B}{\partial m} \dot{v}_x \\ \dot{v}_z = 0 \end{cases}$$

$\frac{\partial E_x}{\partial x} \cdot v_x$
 \uparrow
 $\sin \omega t$
 \uparrow
 $\sim v \sin \omega t$
 \uparrow
 $\sim v^2 \sin^2 \omega t$

$$\Rightarrow \begin{cases} \ddot{v}_x = -\left(\frac{\partial B}{\partial m}\right)^2 v_x + \left(\frac{\partial B}{\partial m}\right) \frac{\dot{E}_x}{B} = -\omega^2 v_x \pm \omega \frac{\dot{E}_x}{B} v_y \\ \ddot{v}_y = -\left(\frac{\partial B}{\partial m}\right)^2 v_y - \left(\frac{\partial B}{\partial m}\right) \left(\frac{\partial B}{\partial m}\right) \frac{E_x}{B} = -\omega^2 v_y - \omega^2 \frac{E_x}{B} \end{cases}$$

* Assuming that the electric field is weak.

'undisturbed orbit' is used to evaluate $E_x(x)$ as an approximation, i.e.,

$$x = x_0 + r_L \sin(\omega t)$$

$$E_x(x) = E_0 \cos(kx) \\ = E_0 \cos[k(x_0 + r_L \sin(\omega t))]$$

$$\Rightarrow \ddot{y} = -\omega^2 y - \omega^2 \frac{E_0}{B} \cos[k(x_0 + r_L \sin \omega t)]$$

→ we would like to find a solution which is the sum of a gyration at ω and a steady drift v_E .
 v_E is what we are interested in. Therefore, gyrotory motion can be taken out by averaging over a cycle.



$$\Delta \begin{cases} \overline{\ddot{v}_x} = 0 = -\omega^2 \overline{v_x} \pm \omega \frac{\overline{E}}{B} \\ \overline{v_x} = \overline{v_0 \sin \omega t} = 0 \\ \overline{E} = \frac{dE}{dx} \cdot \overline{v_x} = \frac{dE}{dx} \overline{v_0 \sin \omega t} = 0 \end{cases}$$

E_x is "periodic" even if it drift along \hat{x} , the average will be zero.

$$\overline{\ddot{y}} = 0 = -\omega_c^2 \overline{y} - \omega_c^2 \frac{\overline{E}}{B} \\ = -\omega_c^2 \overline{y} - \omega_c^2 \frac{E_0}{B} \cos[k(x_0 + r_L \sin \omega t)]$$

$$\cos[k(x_0 + r_L \sin \omega t)] = \cos[kx_0 + k r_L \sin \omega t] \\ = \cos kx_0 \cos(k r_L \sin \omega t) - \sin kx_0 \sin(k r_L \sin \omega t)$$

(for small Larmor radius, i.e., $kr_L \ll 1$)
 $E \equiv kr_L$ $\cos E = 1 - \frac{1}{2} E^2 + \dots$ (1st order)
 $\sin E = E + \dots$ (2nd order)

$$\approx \cos(kx_0) \left(1 - \frac{1}{2} k^2 r_L^2 \sin^2 \omega t\right) - \sin(kx_0) \cdot k r_L \sin(\omega t)$$

↳ $\frac{1}{2}$ after averaging → 0 after averaging

$$\Rightarrow \overline{y} = -\frac{E_0}{B} \cos(kx_0) \left(1 - \frac{1}{4} k^2 r_L^2\right) = -\frac{E_x(x_0)}{B} \left(1 - \frac{1}{4} k^2 r_L^2\right)$$

$\vec{E} \times \vec{B}$ drift:

homogeneous \vec{E} : $\vec{V}_E = \frac{\vec{E} \times \vec{B}}{B^2}$

inhomogeneous \vec{E} : $\vec{V}_E' = \frac{\vec{E} \times \vec{B}}{B^2} \left(1 - \frac{1}{4} k^2 r_L^2 \right)$

$\vec{V}_E' < \vec{V}_E$ since the particle spends more time in regions of weaker \vec{E} .

* The correction term depends on the 2nd derivative of \vec{E} .

* For an arbitrary variation of \vec{E} :
 $ik \rightarrow \nabla$.

$$\Rightarrow \vec{V}_E' = \left(1 + \frac{1}{4} r_L^2 \nabla^2 \right) \frac{\vec{E} \times \vec{B}}{B^2}$$

↑
finite-Larmor-radius effect.

depend on $r_L \rightarrow$ depends on species
 r_L for ion $>$ r_L for electron.

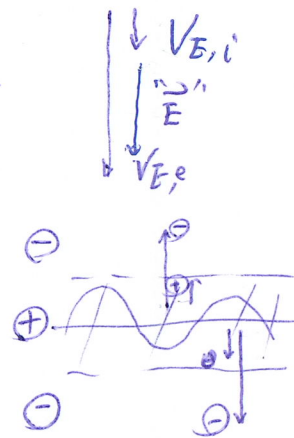
\Rightarrow charge separation occurs

\Rightarrow generating another Electric field.

\Rightarrow drift instability

* Comparing to $V_{\perp B} \propto \vec{B} \times \nabla B \propto k r_L$

$$V_{\perp E} \propto \nabla^2 \propto k^2 r_L^2$$



④ large k , i.e. smaller λ .

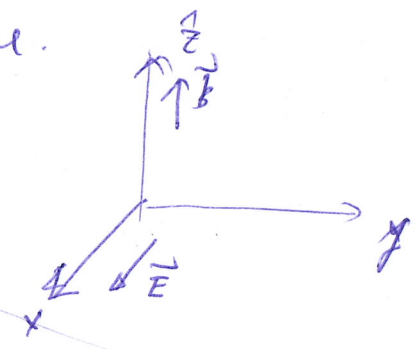
non-uniform- E field effect is more important.

2.4 Time-varying E field

\vec{E} & \vec{B} are uniform in space.
but varying in time.

$$\vec{E} = E_0 e^{i\omega t} \hat{x} \equiv E_x \hat{x}$$

$$\dot{E}_x = i\omega E_0 e^{i\omega t} = i\omega E_x$$



$$m \frac{d\vec{v}}{dt} = q(\vec{E} + \vec{v} \times \vec{B})$$

$$\begin{cases} \dot{v}_x = \frac{qB}{m} v_y + \frac{q}{m} E_x & \Rightarrow \ddot{v}_x = \frac{qB}{m} \dot{v}_y + \frac{q}{m} \dot{E} \\ \dot{v}_y = -\frac{qB}{m} v_x & \Rightarrow \ddot{v}_y = -\frac{qB}{m} \dot{v}_x \end{cases} \quad \omega_c \equiv \frac{|q|B}{m}$$

$$\Rightarrow \ddot{v}_x = -\left(\frac{qB}{m}\right)^2 v_x + \frac{qB}{m} \dot{E} = -\omega_c^2 v_x \pm \omega_c \frac{\dot{E}}{B}$$

$$= -\omega_c^2 v_x \pm i\omega \cdot \omega_c \frac{E_x}{B}$$

note that E_x is oscillating

$$= -\omega_c^2 \left(v_x \mp i \frac{\omega}{\omega_c} \frac{E_x}{B} \right)$$

$$\ddot{v}_y = -\left(\frac{qB}{m}\right)^2 v_y - \left(\frac{qB}{m}\right)^2 \frac{E_x}{B}$$

$$= -\omega_c^2 v_y - \omega_c^2 \frac{E_x}{B}$$

$$\text{let } \begin{cases} \tilde{v}_p \equiv \pm \frac{i\omega}{\omega_c} \frac{E_x}{B} \\ \tilde{v}_E \equiv -\frac{E_x}{B} \end{cases} \Rightarrow \begin{cases} \ddot{v}_x = -\omega_c^2 (v_x - \tilde{v}_p) \\ \ddot{v}_y = -\omega_c^2 (v_y - \tilde{v}_E) \end{cases}$$

To find a solution which is the sum of a drift and a gyrotory motion.

$$\begin{cases} v_x = v_{\perp} e^{i\omega_c t} + \tilde{v}_p \\ v_y = \pm i v_{\perp} e^{i\omega_c t} + \tilde{v}_E \end{cases}$$

↑
gyro motion

$$\begin{aligned} \dot{V}_x &= i\omega_c V_{\perp} e^{i\omega c t} + \dot{\tilde{V}}_p = i\omega_c V_{\perp} e^{i\omega c t} + \left(\frac{\pm i\omega}{\omega_c}\right) \frac{\dot{E}_x}{B} \\ &= i\omega_c V_{\perp} e^{i\omega c t} \pm \frac{i\omega}{\omega_c} \left(\frac{i\omega E_x}{B}\right) \\ &= i\omega_c V_{\perp} e^{i\omega c t} + i\omega \tilde{V}_p \end{aligned}$$

$$\ddot{V}_x = -\omega_c^2 V_{\perp} e^{i\omega c t} - \omega^2 \tilde{V}_p \quad \neq -\omega_c^2 (V_{\perp} - \tilde{V}_p)$$

Note $V_x = V_{\perp} e^{i\omega c t} + \tilde{V}_p$ \uparrow

$$\begin{aligned} \ddot{V}_x &= -\omega_c^2 (V_x - \tilde{V}_p) - \omega^2 \tilde{V}_p \\ &= -\omega_c^2 V_x + (\omega_c^2 - \omega^2) \tilde{V}_p \end{aligned}$$

$$\begin{aligned} \dot{V}_y &= \pm i V_{\perp} (i\omega_c) e^{i\omega c t} + \left(-\frac{\dot{E}_x}{B}\right) = \mp \omega_c V_{\perp} e^{i\omega c t} - \frac{i\omega E_x}{B} \\ &= \mp \omega_c V_{\perp} e^{i\omega c t} + i\omega \tilde{V}_E \end{aligned}$$

$$\begin{aligned} \ddot{V}_y &= \mp i\omega_c^2 V_{\perp} e^{i\omega c t} - \omega^2 \tilde{V}_E \quad \text{Note } V_y = \pm i V_{\perp} e^{i\omega c t} + \tilde{V}_E \\ &= -\omega_c^2 (V_y - \tilde{V}_E) - \omega^2 \tilde{V}_E \\ &= -\omega_c^2 V_y + (\omega_c^2 - \omega^2) \tilde{V}_E \end{aligned}$$

$$\begin{cases} \ddot{V}_x = -\omega_c^2 V_x + (\omega_c^2 - \omega^2) \tilde{V}_p \\ \ddot{V}_y = -\omega_c^2 V_y + (\omega_c^2 - \omega^2) \tilde{V}_E \end{cases} \Leftrightarrow \begin{cases} \ddot{V}_x = -\omega_c^2 + \omega^2 \tilde{V}_p \\ \ddot{V}_y = -\omega_c^2 + \omega^2 \tilde{V}_E \end{cases}$$

extra term \uparrow

\Rightarrow Assuming that E varies slowly, i.e. $\omega^2 \ll \omega_c^2$
 There are two drifting of the guiding center.

① \hat{y} : $\vec{V}_E \perp \vec{B}, \vec{E}$, usual $\vec{E} \times \vec{B}$ drift
 the difference is that it oscillates slowly at the frequency ω .

② \hat{x} : Polarization drift along the direction of \vec{E} .

$$\vec{V}_p = \pm \frac{i\omega}{\omega_c} \frac{E_x}{B}, \quad i\omega \rightarrow \frac{d}{dt}$$

$$\vec{V}_p \equiv \pm \frac{1}{\omega_c B} \frac{d\vec{E}}{dt}$$

\therefore ions & electrons drift in opposite directions:
 \Rightarrow polarization current.

for $Z=1$

$$\begin{aligned} \vec{j}_p &= n \cdot e \cdot (V_{ip} - V_{ep}) \\ &= n \cdot e \cdot \left(\frac{1}{\omega_{ci} B} + \frac{1}{\omega_{ce} B} \right) \frac{d\vec{E}}{dt} \\ &= n \cdot e \cdot \left(\frac{M}{eB^2} + \frac{m}{eB^2} \right) \frac{d\vec{E}}{dt} \\ &= \frac{n(M+m)}{B^2} \frac{d\vec{E}}{dt} \end{aligned}$$

$$= \frac{\rho}{B^2} \frac{d\vec{E}}{dt} \quad \rho: \text{mass density}$$

* $t=0$, ions @ rest, $\vec{B} = B \hat{z}$, $\vec{E} = 0$

$\Rightarrow \vec{E} = E_x(t) \hat{x} \uparrow \Rightarrow$ ion is accelerated $\Rightarrow \underline{V_x > 0}$
 $\Rightarrow \vec{F} = q(\vec{v} \times \vec{B}) \Rightarrow \vec{E} \times \vec{B}$ drift $\underline{\text{drift along } \vec{E}(t)}$
 \Rightarrow If \vec{E} is reversed \Rightarrow decelerates $\Rightarrow \underline{V_x < 0}$
 $\underline{\text{drift along } \vec{E}(t)}$

V_p is a startup drift due to inertia and p 43
 occurs only in the first half-cycle of each
 gyration during which \vec{E} changes

$$V_e \rightarrow 0 \quad \text{with} \quad \frac{\omega}{\omega_c}$$

* an oscillating current \dot{j}_p results from the lag
 due to the ion inertia.

Q 2.5 Time-Varying B field.

* A magnetic field itself cannot impart energy
 to a charged particle.

$\therefore \nabla \times \vec{E} = -\dot{\vec{B}} \Rightarrow$ the \vec{E} associated w/ \vec{B}
 can accelerate particles.

Let $\vec{V}_\perp = \frac{d\vec{l}}{dt} \rightarrow$ particle trajectory

transverse velocity $m \frac{d\vec{V}}{dt} = q(\vec{E} + \vec{V} \times \vec{B}) \cdot \vec{V}_\perp$
 $V_{||}$ is neglected.

$$\Rightarrow \frac{d}{dt} \left(\frac{1}{2} m V_\perp^2 \right) = q \vec{E} \cdot \vec{V}_\perp = q \vec{E} \cdot \frac{d\vec{l}}{dt}$$

$$\begin{aligned} & (\vec{V} \times \vec{B}) \cdot \vec{V}_\perp \\ &= \underbrace{(\vec{V}_\perp \times \vec{B})}_{\perp \vec{V}_\perp} \cdot \vec{V}_\perp = 0 \end{aligned}$$

Integrate over one period.

$$\Delta \left(\frac{1}{2} m V_\perp^2 \right) = \int_0^{2\pi/\omega_c} q \vec{E} \cdot \frac{d\vec{l}}{dt} dt = \oint q \vec{E} \cdot d\vec{l}$$

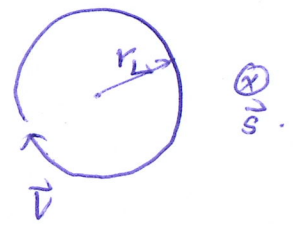
$$= q \int_S (\nabla \times \vec{E}) \cdot d\vec{S} = -q \int_S \dot{\vec{B}} \cdot d\vec{S}$$



Plasma is diamagnetic.

$$\vec{B} \cdot d\vec{s} < 0 \text{ for ion}$$

$$\vec{B} \cdot d\vec{s} > 0 \text{ for electron}$$



$$\begin{aligned} \delta\left(\frac{1}{2} m v_L^2\right) &= \pm |q| \dot{B} \pi r_L^2 \\ &= \pm |q| \dot{B} \pi \left(\frac{v_L}{\omega_c}\right)^2 \\ &= \pm |q| \dot{B} \pi \frac{v_L^2}{\omega_c} \cdot \frac{m}{\pm |q| B} \cdot \frac{2}{2} \\ &= \frac{\frac{1}{2} m v_L^2}{B} \cdot \frac{2\pi \dot{B}}{\omega_c} \\ &= \mu \cdot \frac{\dot{B}}{f_c} = \mu \cdot \frac{f_c \delta B}{f_c} \\ &= \mu \cdot \delta B \end{aligned}$$

$$\dot{B} = \frac{dB}{dt} = \frac{\delta B}{\delta t} \approx f_c \delta B$$

$$\therefore \frac{1}{2} m v_L^2 = \mu \cdot B$$

$$\begin{aligned} \therefore \delta(\mu B) &= \mu \delta B \Rightarrow \delta \mu \cdot B = 0 \\ \text{"} & \text{"} \\ \delta \mu \cdot B + \mu \cdot \delta B & \qquad \qquad \delta \mu = 0 \end{aligned}$$

The magnetic moment is invariant in slowly varying magnetic field.

$$\Rightarrow \frac{\frac{1}{2} m v_L^2}{B} \approx \text{const.} \quad B \uparrow \Rightarrow v_L \uparrow$$

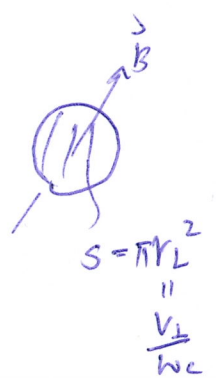
B varies \Rightarrow Larmor orbits expand and contract.

\Rightarrow particle loses / gains transverse energy

Magnetic Flux:

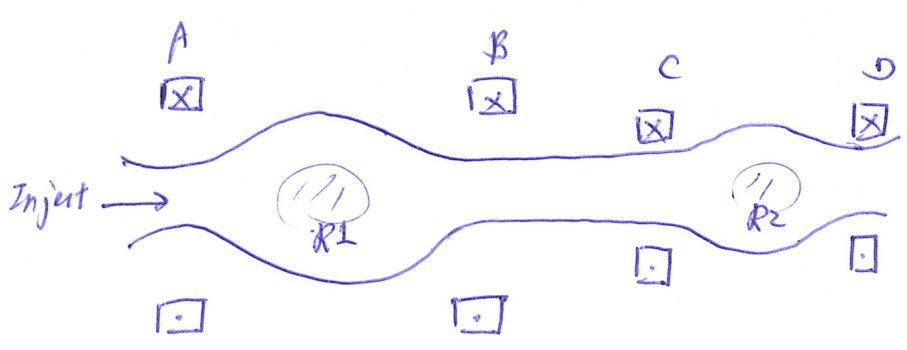
$$\Phi = B \cdot S = B \pi r_L^2 = B \pi \left(\frac{v_L}{\omega_c}\right)^2 = B \pi \frac{v_L^2 \cdot m^2}{q^2 B^2} \frac{2}{2}$$

$= \frac{2\pi m}{q^2} \mu \Rightarrow$ The magnetic flux through a Larmor orbit is constant.



$$B \uparrow \Rightarrow v_{\perp} \uparrow \Rightarrow E_k(T) \uparrow$$

\Rightarrow Adiabatic compression.



- ① Inject into R1
- ② $B_A, B_B \uparrow \Rightarrow$ compression \Rightarrow heating
- ③ $B_A \uparrow \Rightarrow$ push the heated plasma to R2
- ④ $B_C, B_D \uparrow \Rightarrow$ further compression \Rightarrow further heating

276 Summary of guiding center drifts:
 Electric field: V_E

§ 2.6 Summary of guiding center drifts. P46

Electric field : $\vec{V}_E = \frac{\vec{E} \times \vec{B}}{B^2}$ ($\vec{E} \times \vec{B}$ drift)

General force \vec{F} : $\vec{V}_F = \frac{1}{q} \frac{\vec{F} \times \vec{B}}{B^2}$

Gravitational field : $\vec{V}_g = \frac{m}{q} \frac{\vec{g} \times \vec{B}}{B^2}$

Non uniform \vec{E} : $\vec{V}_E = \left(1 + \frac{1}{4} r_L^2 \nabla^2\right) \frac{\vec{E} \times \vec{B}}{B^2}$

Grad-B drift : $\vec{V}_{\nabla B} = \pm \frac{1}{2} v_{\perp} r_L \frac{\vec{B} \times \nabla B}{B^2}$

Curvature drift : $\vec{V}_R = \frac{m v_{\parallel}^2}{q} \frac{\vec{R}_c \times \vec{B}}{R_c^2 B^2}$

Curved vacuum field : $\vec{V}_R + \vec{V}_{\nabla B} = \frac{m}{q} \left(v_{\parallel}^2 + \frac{1}{2} v_{\perp}^2 \right) \frac{\vec{R}_c \times \vec{B}}{R_c^2 B^2}$

Polarization drift : $\vec{V}_p = \pm \frac{1}{\omega_c B} \frac{d\vec{E}}{dt}$

§ 2.7 Adiabatic invariants.

P47

- Action integral $\oint p dq$. p : generalized momentum
 q : generalized coordinate.

In classical mechanics, whenever a system has a periodic motion, the action integral $\oint p dq$ taken over a period is a constant of the motion.

- Adiabatic invariant: If a slow change is made in the system, so that the motion is not quite periodic, the constant of the motion does not change.

- slow \rightarrow compared w/ the period of motion.
 $\rightarrow \oint p dq$ is no longer strictly over a closed path.

§ 2.7.1 The first Adiabatic Invariant. μ .

$$\mu \equiv \frac{m v_{\perp}^2}{2B}$$

The periodic motion involved is Larmor gyration.

$$\int p dq: \quad p: \text{angular momentum. } m v_{\perp} r$$

$$q: \text{coordinate } \theta$$

$$\oint p dq = \int m v_{\perp} r_{\perp} d\theta = 2\pi r_{\perp} m v_{\perp} = 2\pi \frac{m v_{\perp}^2}{\omega_c} = 4\pi \frac{m}{|B|} \mu.$$

$(\omega_c \equiv \frac{|B|}{m})$

μ is a const. of the motion as long as $\frac{m}{|B|}$ is not changed.

Previously, we proved the invariance of μ w/ p48.

$$\omega/\omega_c \ll 1.$$

where ω is the frequency of the rate of change of B

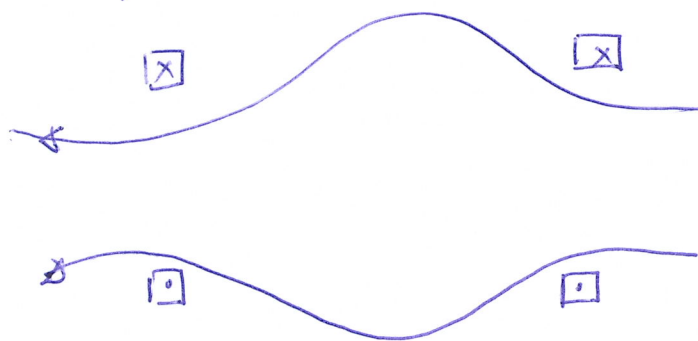
In fact, it holds for $\omega \lesssim \omega_c$.

i.e. μ remains much more nearly constant than B does during one period of gyration.

* Adiabatic invariance of μ is VIOLATED when ω is not small compared w/ ω_c .

Examples:

Magnetic pumping



$$\mu = \frac{m v_{\perp}^2}{2B}$$

If $B \uparrow$ slowly $\rightarrow v_{\perp}^2 \uparrow$ increase slowly

$B \sim$ $\rightarrow v_{\perp}^2 \sim \Rightarrow$ No energy gain

w/ collisions: $v_{\perp}^2 \rightarrow v_{\parallel}^2$ transfer energy from $\perp \rightarrow \parallel$

$\Rightarrow B \uparrow \rightarrow v_{\perp}^2 \uparrow \rightarrow \begin{cases} v_{\parallel}^2 \uparrow \\ v_{\perp}^2 \downarrow \end{cases}$

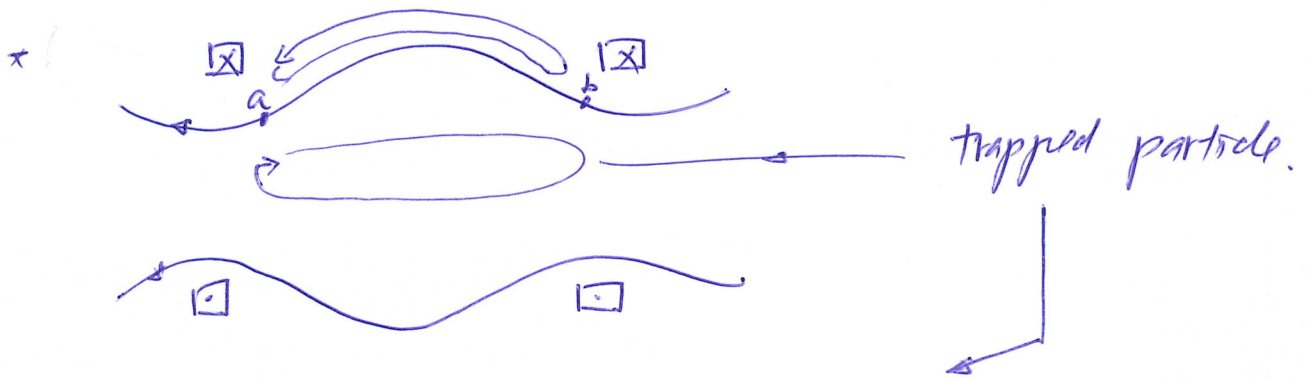
$B \downarrow \rightarrow \begin{cases} v_{\perp}^2 \downarrow \\ v_{\parallel}^2 \text{ doesn't change} \end{cases}$

energy is transferred from $B \rightarrow v_{\perp} \rightarrow v_{\parallel}$

\therefore invariant μ is VIOLATED

\Rightarrow plasma can be heated due to collisions

§ 2.7.2 The second Adiabatic Invariant J, P49



A periodic motion at the "bounce frequency."

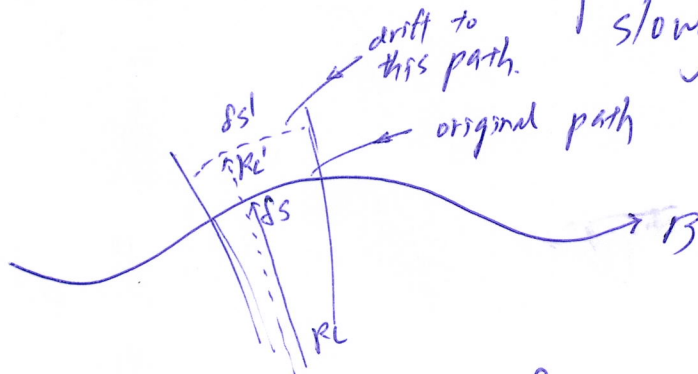
⇒ Action:
$$\oint \underbrace{m v_{||}}_p ds = \oint p dq.$$

* The guiding center drifts across field lines, the motion is not exactly periodic. The constant of the motion becomes an adiabatic invariant:

Longitudinal invariant J.

$$J = \int_a^b v_{||} ds.$$
 ← half cycle between two turning points a, b.

J is invariant in $\left\{ \begin{array}{l} \text{a static, nonuniform } B \\ \text{slowly time-varying } B. \end{array} \right.$



Consider $v_{||} SS$ first. SS : a segment of the path along \mathbf{k} due to the drift, particle drifts to SS' after Δt .

show first.

$$\frac{ss'}{R_c} = \frac{ss'}{R_c'}$$

$$\Rightarrow \frac{ss'}{ss} = \frac{R_c'}{R_c} \Rightarrow \frac{ss'}{ss} - 1 = \frac{R_c'}{R_c} - 1$$

$$\Rightarrow \frac{\frac{ss' - ss}{\Delta t}}{ss} = \frac{R_c' - R_c}{\Delta t R_c} \rightarrow V_{gc} \text{ in } \hat{R} \text{ is } \frac{R_c' - R_c}{\Delta t}$$

$$\vec{V}_{gc} \cdot \hat{R}_c = \frac{R_c' - R_c}{\Delta t}$$

$$\vec{V}_{gc} \cdot \frac{\vec{R}_c}{R_c}$$

Note that $\vec{V}_{gc} = \vec{V}_{\perp B} + \vec{V}_R = \pm \frac{1}{2} v_{\perp} r_L \frac{\vec{B} \times \vec{\nabla} B}{B^2} + \frac{mv_{\parallel}^2}{8} \frac{\vec{R}_c \times \vec{B}}{R_c^2 B^2}$

Grad-B drift Curvature drift

$$\frac{1}{ss} \frac{dss}{dt} = \vec{V}_{gc} \cdot \frac{\vec{R}_c}{R_c^2} = \underbrace{\left(\pm \frac{1}{2} v_{\perp} r_L \right)}_{\frac{mv_{\perp}}{2B}} \frac{\vec{B} \times \vec{\nabla} B}{B^2} \cdot \frac{\vec{R}_c}{R_c^2}$$

\uparrow
 $\perp \vec{R}_c$
 $\Rightarrow \vec{R}_c = i$

$$= \frac{1}{2} \frac{mv_{\perp}^2}{2B^3} (\vec{B} \times \vec{\nabla} B) \cdot \frac{\vec{R}_c}{R_c^2} \rightarrow \text{the rate of change of } ss \text{ as seen by the particle.}$$

Total energy:

$$W = \frac{1}{2} m v_{\parallel}^2 + \frac{1}{2} m v_{\perp}^2 = \frac{1}{2} m v_{\parallel}^2 + \mu B \equiv W_{\parallel} + W_{\perp}$$

\Rightarrow to get v_{\parallel} -----

$$\Rightarrow v_{\parallel} = \sqrt{\frac{2}{m} (W - \mu B)}$$

Note that W, μ are invariant only B changes.

$$\frac{1}{v_{\parallel}} \frac{dv_{\parallel}}{dt} = \frac{1}{\sqrt{\frac{2}{m} (W - \mu B)}} \times \frac{1}{2} \sqrt{\frac{2}{m}} \frac{-\mu}{\sqrt{W - \mu B}} \frac{dB}{dt} = -\frac{1}{2} \frac{\mu \dot{B}}{W - \mu B} = -\frac{1}{2} \frac{\mu \dot{B}}{W_{\parallel}}$$

$$= -\frac{\mu \dot{B}}{m v_{\parallel}^2}$$

$$\dot{\vec{B}} = \frac{d\vec{B}}{dt} = \frac{\partial \vec{B}}{\partial t} + \underbrace{\frac{\partial \vec{B}}{\partial \vec{r}} \cdot \frac{d\vec{r}}{dt}}_{\text{change of } \vec{B} \text{ seen by the particle due to the guiding center motion.}}$$

\therefore static \vec{B} \rightarrow change of \vec{B} seen by the particle due to the guiding center motion.

$$= \vec{V}_{gc} \cdot \nabla \vec{B} = \frac{mV_{\perp}^2}{\hbar} \frac{(\vec{R}_c \times \vec{B})}{R_c^2 B^2} \cdot \nabla \vec{B}$$

$$\Rightarrow \frac{\dot{V}_{\parallel}}{V_{\parallel}} = -\frac{\mu}{\hbar} \frac{(\vec{R}_c \times \vec{B})}{R_c^2 B^2} \cdot \nabla \vec{B} = -\frac{1}{2} \frac{m}{\hbar} \frac{V_{\perp}^2}{B} \frac{(\vec{B} \times \nabla \vec{B}) \cdot \vec{R}_c}{R_c^2 B^2}$$

The fractional change in V_{\parallel} is

$$\frac{1}{V_{\parallel} ds} \cdot \frac{d}{dt} (V_{\parallel} ds) = \frac{1}{ds} \frac{ds}{dt} + \frac{1}{V_{\parallel}} \frac{dV_{\parallel}}{dt}$$

} Show first. \star

$$= \frac{1}{2} \frac{mV_{\perp}^2}{\hbar B^3} (\vec{B} \times \nabla \vec{B}) \cdot \frac{\vec{R}_c}{R_c^2} - \frac{1}{2} \frac{m}{\hbar} \frac{V_{\perp}^2}{B} \frac{(\vec{B} \times \nabla \vec{B}) \cdot \vec{R}_c}{R_c^2 B^2}$$

$$= 0$$

$$\Rightarrow V_{\parallel} ds = \text{const.}$$

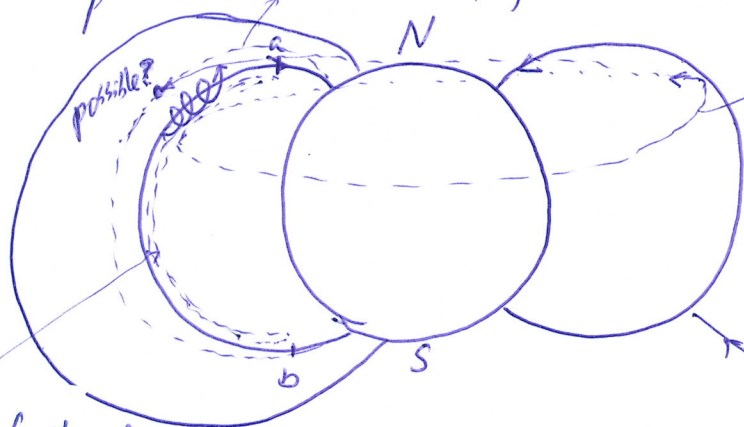
\therefore At turning point, $V_{\parallel} \sim 0$.

$$\therefore J = \int_a^b V_{\parallel} ds = \text{const.}$$

Example:

another field line at the same μ longitude but different altitude

slowly drift in longitude around the earth



bounce back & forth due to magnetic mirror.

$\therefore \mu$ is invariant.

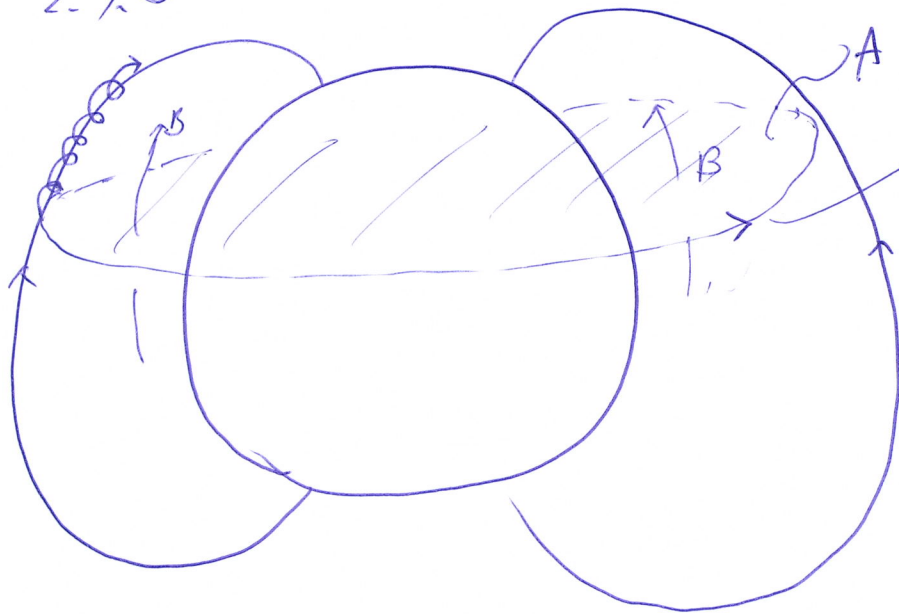
$\therefore |\vec{B}|$ remains the same at the turning point.

actual field is distorted by such effects as the solar wind.

$J = \int_a^b v_{\perp} ds$ invariant \because J determines the length^{PS:}
 of the line of force between turning points, and
 no two lines have the same length between
 points with the same $|B|$

\Rightarrow the particle returns to the same line of force
 even in a slightly asymmetric field.

7. 2.2.3 The third Adiabatic Invariant $\bar{\Phi}$



slow drift of
 a guiding center
 around the earth
 \rightarrow periodic motion

$\bar{\Phi} \equiv \int \vec{B} \cdot d\vec{A}$: total magnetic flux enclosed by
 the drift surface.

\Rightarrow the particle will stay on a surface such that
 the total number of lines of force enclosed
 remains const.

\Rightarrow Few applications because most fluctuations of B
 occur on a time scale short compared w/ the
 drift period.